# Geophysical Trends inferred from 20 years of AIRS infrared global observations

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# Key Points:

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8	•	The 20+ year low noise high spectral radiance record by NASA's Atmospheric In-
9		frared Sounder contains detailed vertical information about surface and atmospheric
10		temperature and water vapor.
11	•	Trends from the radiance maeasurements are analyzed in a novel way for long-term
12		climate studies, different than traditional use of infrared radiaces in daily retrievals

climate studies, different than traditional use of infrared radiaces in daily retrievals or assimilation into Numerical Weather Prediction Models.

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# 14 Abstract

NASA's Atmospheric Infrared Sounder has been in near-continuous operation since 15 September 2002. The  $\sim 3$  million daily spectral observations contain detailed informa-16 tion about surface and atmospheric temperature, water vapor and trace gases such as 17  $CO_2$  and  $CH_4$ , as well as clouds and aerosols. In this paper we obtain climate thermo-18 dynamic trends using 20 years of AIRS observations by working exclusively with the trends 19 observed in the AIRS radiance time series. This is achieved by first binning the observed 20 spectra into nominal  $3 \times 5$  degree latitude/longitude spatial subsets using 16 day inter-21 22 vals, after which a quantile-based algorithm selects nominally clear scenes for each grid box in order to construct the clear scene radiance spectrum time series. De-seasonalized 23 spectral anomalies and spectral trends are then obtained from the time series, which are 24 converted into geophysical trends using a physical retrieval for each grid box. This approach 25 is completely different from traditional operational use of infrared data for trending, whereby 26 anomalies/trends are generated either after daily retrievals, or after assimilation into NWP 27 models. Our approach rigorously ties the derived geophysical trends to the observed ra-28 diance trends, and requires orders-of-magnitude fewer computational resources and time 29 than re-analysis or traditional Level 2 retrievals. The retrieved trends are compared to 30 trends derived from four other products : ERA5, MERRA2 reanalysis model fields and 31 the NASA Level3 AIRS v7 and NASA Level 3 CLIMCAPS v2. Our retrieved surface tem-32 perature trends agree quite well with ERA5 re-analysis, CLIMCAPS L3 and the GISS 33 surface climatology trends. Atmospheric temperature profile trends exhibit some vari-34 ability amongst all these data sets, especially in the polar stratosphere. Water vapor pro-35 file trends are nominally similar amongst all data sets except for the AIRS v7 which ex-36 hibits trends with a different sign in the mid troposphere. Note that infrared sounders 37 lose water vapor sensitivity close to the surface making intercomparisons of column water 38 trends problematic. Spectral closure between observation trends versus those computed 39 by running all the NWP re-analysis and official NASA L3 monthly fields though a (clear 40 sky) radiative transfer code is discussed, with the major differences arising in the wa-41 ter vapor sounding region. 42

# 43 Plain Language Summary

The new generation of infrared sounders, designed for weather forecasting purposes, 44 have been in orbit around the Earth for a long enough time to enable anomaly and trending 45 studies for climate purposes. Traditionally their daily obtained radiance data has been 46 used for operational atmospheric state retrievals, or assimilation into Numerical Weather 47 Prediction models, after which climate anomaly studies are made. In this paper we use 48 the raw radiance spectral data to form radiance anomalies and trends, after which we 49 do a one step atmospheric state retrieval. This novel approach has the benefit of using 50 only stable channels together with easily understood assumptions and well tested retrieval 51 algorithms to do the trend or anomaly geophysical retrieval, which has full error characterization. 52 53

# 54 1 Introduction

NASA's Atmospheric Infrared Sounder (AIRS) became operational in September 55 2002, as the first of the new generation of low noise, high stability hyperspectral sounders, 56 making Top of Atmosphere (TOA) radiance observations at a typical 15km (at nadir) 57 horizontal resolution. Follow on instruments with similar characteristics and abilities in-58 clude Eumetsat's Infrared Atmospheric Sounding Interferometer (IASI) and NOAA's Cross 59 Track Infrared Sounder (CrIS), operational since June 2007 and March 2012 respectively. 60 The latter two already have follow on missions planned till the 2040s, and together these 61 three sounders will provide scientists with a 40 year high quality, near continuous ob-62 servational dataset for climate anomaly and trending studies. 63

Infrared radiances contain a wealth of information, including but not limited to sur-64 face temperature, atmospheric temperature and water amount, and mixing ratios of green-65 house gases such as carbon dioxide  $CO_2$ ,  $CH_4$  and  $N_2O$ . Measurements by visible imagers 66 which have  $\sim 1$  km horizontal resolution or better King et al. (2013) suggest global cloud 67 free fractions of  $\sim 30\%$ , but the 15 km footprint of typical sounders means at most 5% 68 of the hyperspectral observations can be considered "cloud-free." Current operational NASA 69 L2 products come from cloud clearing the observed radiances, which introduces errors 70 and use the method of cloud clearing on observed radiances in partly cloudy scene conditions 71 before doing the geophysical retrieval. The cloud clearing method solves for an estimate 72 of clear column radiances by examining adjacent Fields of View (FOVs) to estimate the 73 cloud effects on observed allsky radiances, assuming any differences are solely due to different 74 cloud amounts in each FOV, and significantly increases geophysical retrieval yields (to 75 about 50-60%) Smith and Barnet (2023). This does introduce increased noise in the cloud 76 cleared radiances of the lower atmosphere sounding channels; in addition the subsequent 77 retrieval depends on the first guess (which is a neural net for AIRS v7 and MERRA2 re-78 analysis for CLIMCAPS v2). The The reader is referred to Susskind et al. (2003); Smith and Barnet (2020) 79 Susskind et al. (2003); Smith and Barnet (2020, 2023) for more details. 80

In this paper we work directly in radiance space and form either anomalies or trends 81 from the underlying well characterized and understood radiances Strow and DeSouza-82 Machado (2020), in order to do a geophysical trend or anomaly retrieval. The work pre-83 sented here, once the averaged/sorted data is available, can be processed in hours to days, 84 and can be duplicated by small research groups with ease. Moreover, our novel approach 85 has zero temperature *a-priori* and minimal water vapor *a-priori*. This completely sidesteps 86 time variability and the accuracy of the *a-priori* which causes errors in the retrievals, 87 and ensures our work examines trends directly inferred from the radiances versus those 88 from traditional methods, leading , This leads to more unbiased results that directly high-89 light the conditions (for example stratospheric water vapor) where the sensor has lim-90 ited sensitivity. 91

The approaches used in this work are therefore very different than climate anoma-92 lies or trends from reanalysis products or traditional Level 2 retrievals, neither of which 93 are tailored for climate trends. Reanalysis uses a wide range of observations and are only 94 ereated within very large organizations, and represent the most commonly used climate 95 data sets. They products assimilate individual sensor scenes from many different instruments. 96 and may have discontinuities as different instruments come online or go offline. Tradi-97 tional Level 2 (and Level 3 products derived from Level 2) retrieve the atmospheric state 98 for individual scenes (or effective cloud-cleared radiance derived from a 3x3 grid of in-99 dividual scenes). Both reanalysis and Level 2 products require large computational re-100 sources, that preclude full dataset re-processing to help fully understand trends. A main 101 characteristic of traditional L2 retrievals is the requirement for a good *a-priori* state for 102 each inversion, making errors in the *a-priori* difficult to distinguish from true variabil-103 ity in the data, especially with regard to trends. 104

The stability and accuracy of the AIRS instrument is documented in recent work 105 on analyzing 16 years of AIRS radiance anomalies over cloud-free ocean Strow and DeSouza-106 Machado (2020). Geophysical retrievals on the anomalies yielded CO<sub>2</sub>, CH<sub>4</sub>, N<sub>2</sub>O and 107 surface temperature time series that compared well against in-situ data from NOAA Global 108 Monitoring Laboratories (GML) trace gas measurements and NOAA Goddard Institute 109 of Space Studies (GISS) surface temperature data respectively. A significant difference 110 between this paper and Strow and DeSouza-Machado (2020) is the nominally clear scenes 111 used in this paper are selected uniformly from all over the Earth, while the clear scenes 112 in the latter were zonal averages which were sometimes concentrated in certain regions. 113

In this paper we expand upon our initial zonal clear sky analysis, to derive geophysical trends from 20 years (September 2002 - August 2022) of AIRS measurements over  $\sim 3 \times 5$  degree tiles covering the Earth, chosen such that the number of observations

in each tile is roughly equal. An important concept introduced is spectral closure, whereby 117 the observed clear sky spectral radiance trends are compared to spectral trends produced 118 by running the monthly reanalysis or official NASA retrieved AIRS L3 products through 119 an accurate clear sky radiative transfer code; close agreement in different sounding re-120 gions (such as 640-800 cm<sup>-1</sup> for temperature and CO<sub>2</sub>, 1350-1640 cm<sup>-1</sup> for water va-121 por, 1000-1150 cm<sup>-1</sup> for  $O_3$ ) between the computed and actual observed spectral trends 122 imply that trends from those geophysical parameters used in the computations are re-123 alistic while disagreement suggests otherwise. A companion paper will utilize the geo-124 physical trend results to derive Outgoing Longwave Radiation (OLR) trends and non-125 local clearsky feedback parameters. Nominally clear scenes for each tile are picked out 126 using a quantile approach; from the time series, radiances trends are made over the en-127 tire Earth, from which geophysical trends are retrieved. 128

Observed infrared spectral trends from AIRS has already been a focus of earlier 129 work by Huang et al. (2023) who studied a slightly shorter time period (2002-2020) while 130 Raghuraman et al. (2023) converted the radiances to Outgoing Longwave radiation (OLR), 131 but neither study involve retrievals from spectral trends to geophysical trends. Instead 132 they convert various model trends (such as ERA5) to spectral trends and compare against 133 the observed spectral trends. Our earlier work shows we can accurately account for the 134 effects of GHG forcings (Strow et al., 2021). In this paper we remove these GHG forc-135 ings from the observed AIRS spectral trends to concentrate on atmospheric temperature 136 and water vapor and surface temperature, while the papers by Huang et al. (2023); Raghu-137 raman et al. (2023) include the GHG forcings in the model generated spectral trends. 138 Another noteworthy examination of the time evolution of high spectral resolution infrared 139 radiances (converted to spectral outgoing longwave radiation (OLR) fluxes) by (Whitburn 140 et al., 2021) covered 10 years (2007-2017) of IASI observations. They confirmed that the 141 IASI-derived fluxes agreed well with increases in GHG gas concentrations and El-Nino 142 Southern Oscillation (ENSO) events within that time frame. A more recent paper Roemer 143 et al. (2023) used the 10 year IASI data to derive allsky longwave feedback spectral com-144 ponents (water vapor, CO<sub>2</sub>, window, ozone) and total values, while also estimating clear 145 sky feedback values. 146

We will refer to our results as the AIRS Radiance Trends (AIRS RT). Compar-147 isons are made against monthly output from the European Center for Medium Weather 148 Forecast fifth generation reanalysis (ERA5) Hersbach et al. (2020) and NASA's second 149 generation Modern-Era Retrospective analysis for Research and Applications (MERRA2) 150 Gelaro and Coauthors (2017), and also against the official monthly AIRS L3 products 151 which are AIRS v7 L3 Susskind et al. (2014); Tian et al. (2020) and CLIMCAPS v2 L3 152 Smith and Barnet (2019, 2020). Detailed geophysical trends and spectral closure stud-153 ies are presented for the ascending (daytime (D)), descending (nightime (N)) and D/N154 averages. 155

# <sup>156</sup> 2 Datasets used in this study

Three main types of datasets are used in this study. The first is the AIRS L1C ra-157 diance dataset we analyzed for this paper, which has both daytime (D) and nightime (N) 158 (ascending and descending) views of the planet. Second is the monthly operational L3 159 retrieval data, which are the AIRS v7 and the CLIMCAPS v2 products, also separated 160 into D/N data. Finally we also compared to trends from ERA5 and MERRA2 monthly 161 reanalysis model fields. The ERA5 monthly dataset is available in 8 averaged time steps. 162 so we match to the average AIRS overpass times and compute (D/N) data over the 20 163 years, while MERRA2 monthly model fields are only available as one time step; included 164 here for completeness we mention the NASA GISS surface temperature dataset, which 165 like MERRA2 is only available as one set per month monthly mean. This means four 166 of the datasets : AIRS RT (from AIRS L1C), AIRS L3 and CLIMCAPS L3, and ERA5 167 are separable into D/N, while the other two (MERRA2 and GISS) are only available as 168

a diurnal averaged value. We describe these datasets in more detail below. In addition
 we also briefly mention other datasets that we use.

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# 2.1 The AIRS instrument and L1C dataset

The Atmospheric Infrared Sounder (AIRS) on board NASA's polar orbiting EOS/Aqua 172 platform has 2378 channels, covering the Thermal Infrared (TIR) spectral range (roughly 173 649-1613 cm<sup>-1</sup> ) and shortwave infrared (2181-2665 cm<sup>-1</sup> ). The full widths at half max-174 imum satisfy  $\nu/\delta\nu \sim 1200$ . The (spectral dependent) noise is typically  $\leq 0.2$ K. The orig-175 inal L1b radiance dataset suffers from spectral gaps and noise contaminated data as de-176 tectors slowly fail. These limitations are addressed using a 2645 L1c channel dataset, where 177 spectral gaps and some of the noise "pops" are filled in using principal component recon-178 struction Manning et al. (2020) and is the dataset used to subset radiances analyzed in 179 this paper. However we note that the results described in this paper used only the ac-180 tual observed radiances in pristine, stable channels described in Strow et al. (2021) and 181 none of the synthetic channels. The Aqua platform is a polar orbiting satellite with 1.30 182 am descending (night time over equator) and 1.30 pm ascending (daytime over equator) 183 tracks. Each orbit takes about 90 minutes, with the 16 passes yielding almost twice daily 184 coverage of the entire planet. About  $\sim 3$  million AIRS spectral observations have been 185 obtained daily since AIRS became operational in late August 2002. The instrument has 186 provided data almost continuously since then though there have been some shutdowns 187 (each spanning a few days) such as during solar flare events. 188

In this paper we use the re-calibrated 2645 channel L1C radiance data Strow and 189 DeSouza-Machado (2020) instead of the 2378 L1B data. 20 years (spanning September 190 1, 2002-August 31, 2022) of AIRS L1C radiance data is gridded into 4608 tiles covering 191 the Earth : 72 longitude boxes which are all  $5^{\circ}$  in width, and 64 latitude boxes which are 192 approximately  $2.5^{\circ}$  in width at the tropics but wider at the poles to keep the number of 193 observations per 16 day intervals (which is the repeat cycle of the AIRS orbit on the Aqua 194 satellite) roughly the same. This way there are  $\sim 12000$  observations per 16 days per tile. 195 which are roughly equally divided between the ascending/daytime (D) and descending/nigh-196 time (N) tracks. In this paper we discuss results for both the ascending and descending 197 tracks using a retrieval based on the longwave (LW) and midwave (MW) regions of the 198 spectrum (640-1620 cm<sup>-1</sup> or 6-15  $\mu m$ ). 199

In this paper our trend retrievals use only the AIRS channels are stable in time, 200 as quantified in Strow et al. (2021). For example the shortwave (SW) channels are drift-201 ing at a higher rate than the LW/MW channels, which can lead to incorrect surface tem-202 perature rates, and are avoided in this paper. Similarly there are many channels in 203 the LW and MW whose detectors are drifting in time, and which are also not used here. 204 For example there are some higher wavenumber (shorter wavelength) channels past the 205 ozone band which have a significant drift in time, possibly due to changes in the polar-206 ization of the scan mirror coating with time. Therefore compared to other AIRS oper-207 ational products used in this paper, our results use channels that are demonstrated to 208 have high stability Strow et al. (2021). We do note that some of the observed drifts in 209 the AIRS channels stabilized after 6 years, so their impact is reduced when looking at 210 20 year trends. 211

In "clear sky" scenes, the window region would be dominated by the effects of water 212 vapor continuum absorption, which is largest for hot and humid tropical scenes and almost 213 negligible for cold, dry scenes. Scattering and absorption by liquid and ice clouds also 214 affects the window region (800-1000 and 1100-1250, and 2400-2800). For each tile, we 215 use the 1231.3 observation as our window channel (AIRS L1C channel ID = 1520), and 216 form the quantiles of the observed Brightness Temperature (BT) for each 16 day observation 217 period. BT 1231 therefore serves as a measure for the cloudiness of an observation : if 218 there are no or low clouds, it will effectively measure the surface temperature, but as the 219

elouds get thicker and higher, it will measure colder temperatures. Quantiles 0.50, 0.80,
0.90, 0.95 and 0.97 were among those chosen; the first would be considered the "median"
observation, containing clear and cloudy scenes. In a subsequent section we show Q0.90
onwards can be considered "almost free of clouds." Our retrievals using this dataset are
referred to as in what follows.

2.2 Reanalysis Model fields

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The ERA5 fifth generation reanalysis product from the European Center for Medium 226 Range Weather Forecasts is freely available on monthly timescales from the Copernicus 227 Climate Data Store. This monthly dataset is output at 37 pressure levels at 0.25° horizontal 228 resolution Hersbach et al. (2020), which is further subdivided into eight 3-hour averages 229 per month (corresponding to 00,03,06,...21 UTC). For each month from September 2002-230 August 2022 we downloaded the surface temperature and pressure fields, as well as at-231 mospheric temperature, water vapor and ozone fields. These are then colocated to each 232 tile center using 2D spatial interpolation, as well as time interpolated according to the 233 average AIRS overpass time as a function of month. From the resulting monthly time-234 series of reanalysis model fields for each tile, we generated (a) thermodynamic trends for 235 surface temperature, air temperature, water vapor and ozone model fields (b) a 20 year 236 average thermodynamic profile in order to produce jacobians for the linear trend retrievals 237 (c) by using the model fields as input to the clear sky SARTA radiative transfer code Strow, 238 Hannon, DeSouza-Machado, et al. (2003) a monthly time series of clear sky radiances 239 for each tile was generated, from which we could compute radiance trends. We did this 240 for both the ascending and descending datasets. 241

The MERRA version 2 (MERRA2) re-analysis used in this paper is the second gen-242 eration Gelaro and Coauthors (2017) product from NASA's Global Modeling and As-243 similation Office. The monthly data we use is available on 42 pressure levels at a hor-244 izontal resolution of  $0.5^{\circ} \times 0.625^{\circ}$ , but only one monthly mean diurnally averaged out-245 put is available per month. Similar to the ERA5 output, we colocated the MERRA2 sur-246 face temperature, atmospheric temperature, water vapor and ozone fields to our tile cen-247 ters for each month starting September 2002 in order to produce a time series of radi-248 ance and model output, from which radiance and thermodynamic trends could be com-249 puted for comparisons against other datasets in this study; similar to above we also gen-250 erated a monthly time series of clear sky radiances for each tile, from which we could com-251 pute clear sky radiance trends based on MERRA2. 252

The NASA Goddard Institute of Space Studies (GISS) surface temperature data v4 surface temperature data 2023 (2005); Lenssen et al. (2019) is a monthly dataset based primarily on near surface temperatures land stations, and data from ships and buoys. As with MERRA2 we obtained one monthly mean dataset per month, which we could not separate into descending (N) or ascending (D) tracks.

2.3 AIRS L3 Products

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NASA routinely produces two retrievals from the AIRS L1C data observed each 259 day, which are AIRS v7 Susskind et al. (2014); Tian et al. (2020) and CLIMCAPS v2 260 Smith and Barnet (2019, 2020). Both use the cloud clearing process but there are sig-261 nificant algorithmic differences; in particular the AIRS v7 product is initialized by a neu-262 ral net, while CLIMCAPS uses MERRA2 for its initialization. The L2 products are then 263 individually turned into L3 monthly products, for both the ascending (daytime) and de-264 scending (nighttime) data. The timeseries of thermodynamic profiles were used as in-265 put to the clear sky SARTA RTA to generate radiances, after which radiance trends and 266 thermodynamic trends are also produced. 267

# 268 2.4 Other L3 Products

The Microwave Limb Sounder (MLS) monthly binned water vapor (H2O) mixing ratio dataset Lambert et al. (2021), which contains data at spatial coverage  $\pm 82^{\circ}$  latitude, at a spatial resolution of  $4^{\circ} \times 5^{\circ}$  and useful vertical range between 316 and 0.00215 hPa was used in this paper to improve retrieval trends in the upper atmosphere.

# <sup>273</sup> 3 Filtering the Observational Data for clear scenes

Here we discuss the "clear scene" selection from all the observed data stored for each of the 72 × 64 tiles. Ideally we would prefer to use a MODIS cloud fraction product (1 km) colocated to the 15 km AIRS footprints, but this is presently unavailable. Our earlier work used an uniform clear flag over ocean Strow et al. (2021) which will not work well over land because of surface inhomogeneity. In this section we discuss an alternative clear filter based on the hottest 10 percent of AIRS observations that are present inside any 16 day tile, over any location.

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# 3.1 Observed BT1231 Distributions

The left panel of radiances measured in thermal infrared window region (800-1000 282  $cm^{-1}$  and 1100-1250  $cm^{-1}$ ) are dominated by the effects of the surface temperature, water 283 vapor continuum absorption and cloud/aerosol effects. The effects of water vapor continuum 284 absorption is largest in hot and humid tropical scenes (depressing the observations relative 285 to surface temperatures by about 5-6 K, which reduces to about 2 K at  $\pm$  50°) and is 286 almost negligible for cold, dry scenes (less than 1 K). Scattering and absorption by liquid 287 and ice clouds also affects the window region (Deep Convective Clouds can depress the 288 window channel observations by as much as 100 K relative to surface temperatures). For 289 each tile, we use the 1231.3  $\mathrm{cm}^{-1}$ observation as our representative window channel (AIRS 290 L1C channel ID = 1520), as it is minimally impacted by weak water vapor lines. Changed 291 to Brightness temperature (BT) the observation in this 1231.3  $\mathrm{cm}^{-1}$ channel (BT1231) 292 therefore serves as a measure for the cloudiness of an observation ; if there are no or low 293 or optically thin clouds, it will effectively measure the surface temperature, but as the 294 clouds get thicker and higher, it will measure the cold cloud top temperatures. For any 295 tile during any 16 day observation periods, we can compute quantiles  $\mathcal{Q}$  based on the 296 observed BT1231 to screen between cloudy and partially clear scenes. We chose different 297 BT1231 quantiles (so quantile Q0.XY will have a numerical value  $BT1231_{Q0,XY}$  associated 298 with it) and show below the data contained between Q0.90 and Q1.00 can be considered 299 "almost free of clouds." 300

Figure 1 shows the zonally averaged histograms for a all the BT1231 observations 301 for a chosen 16 day timestep in the form of a zonally averaged histogram (normalized 302 probability distribution functions (PDFs)), with latitude on the vertical axis and BT1231 303 on the horizontal axis. The colorbar is the PDF value, and we used data spanning Au-304 gust 27, 2012 - September 11, 2012 which is approximately half way through the 20 year 305 AIRS mission dataset used in this paper. The colorbar is the mean histogram (normalized 306 probability distribution functions (PDFs)) using the data in that 16 day time period. From 307 this we plot the curves show the zonally averaged BT1231 values of the minimum ( $\frac{Q0.00}{Q0.00}$ ) 308 (20.00) in dark cyan, mean (thick red), mean, median ((20.5020.50) in orange), maximum 309 (Q1.00Q1.00 in light cyan); also shown are the a handful of other zonally averaged BT1231 310 values<del>of Q0.80, Q0.90</del>, for example Q0.80, Q0.90 (thick black curve), Q0.95 and Q0.97. 311 The BT1231 channel has the lowest expected absorption due to water vapor in the longwave 312 portion of the spectrum, and so is expected to sense the surface temperature unless the 313 scene is cloudy in which case it would be expected to sense the cloud top temperature. 314 In this way the histogram should exhibit the characteristics of the cloud conditions observed 315 in the 16 day period. 20.95 and 20.97. The distributions are skewed to the left (negative 316 skewness), as confirmed by the mean being less than the median. We also point out that 317

even Q0.80 sees much of the surface from the southern tropics to the northern polar region.
 The 220 K horoizontal axis cutoff means we do not see the very cold (190 K) observations
 over the winter Antarctic.

The figure shows the expected qualitative features, for example (1) the tropical PDFs 321 peak at around 295 K, but show some warmer observations, as well much colder obser-322 vations (below 230 K) corresponding to Deep Convective Clouds (DCC); this gives a dy-323 namic range of almost 100 K at the tropics (2) the BT1231 observed over the Southern 324 Polar (polar winter) regions are much colder than the BT1231 observed over the North-325 326 ern Polar (polar summer) regions and (3) the reddish peaks in the  $30^{\circ}N$  -  $40^{\circ}N$  are a combination of the marine boundary layer (MBL) clouds and warmer summer land temper-327 atures. 328

It is evident the distributions are skewed to the left (negative skewness), as confirmed 329 by the mean being less than the median. We also point out that even Q0.80 sees much 330 of the surface from the southern tropics to the northern polar region. The right panel 331 of Figure 1 shows the same information, except presented as a cumulative histogram, with 332 a value of 0 at the hot end (340 K) and 1 at the cold end (180 K); again one sees the Q0.90 333 quantile envelopes the hottest 10% of the observations as expected. The cutoff of 220 334 K in the plots does not allow the plot to extend to show the very cold (190 K) observations 335 over the winter Antarctic. 336

<sup>337</sup> Zonally averaged BT1231 histograms for an 2012/08/27 - 2012/09/11 timespan (colorbar) <sup>338</sup> and quantiles (curves). The thick black curve is the Q0.90 quantile (and above) used in <sup>339</sup> this paper, and is very close to the maximum. The left hand panel shows the normalized <sup>340</sup> histogram (probability distribution function) as a function of latitude and temperature <sup>341</sup> bin; the right hand panel shows the cumulative distribution function, though starting <sup>342</sup> from the hotter side (cdf(340 K) = 0.0, cdf(180 K) = 1.0).

Measurements by visible imagers which have  $\sim 1 \text{ km}$  horizontal resolution or better 343 King et al. (2013) suggest global cloud free fractions of  $\sim 30\%$ , but the 15 km footprint 344 of typical sounders means at most 5% of the hyperspectral observations can be considered 345 "cloud-free." In the tropics, the higher amounts of water vapor means the observed BT1231 346 for a clear scene would be reduced by a 5-6 K due to water vapor continuum (which on 347 average reduces to about 2 K at  $\pm$  50, and 1 K at the polar regions). Figure 1 shows on 348 <del>average the</del> on average the cloud effect at the tropics is an additional modest 20 K (dif-349 ference between  $\frac{Q0.90 \text{ and } Q0.50}{Q0.90}$  and Q0.50 compared to the 100 K dynamic range. 350 This is because the cloud fractions and cloud decks in the individual observations have 351 effectively more clouds (with larger cloud fraction in the FOV) lower in the atmosphere 352 than higher up; the net effect is that in the window region the atmosphere is on aver-353 age radiating from the lower (warmer) altitudes, and so Q0.80 to Q1.00 onwards spectra 354 whose BT1231 values are larger than  $BT1231_{0,0,80}$ , see much of the surface emission as 355 well. 356

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We now use the above plots to select "almost clear" scenes. For any one tile, we define set  $\Psi_{0,XY}$  to have all observations *i* whose BT1231 lies between quantiles  $Q_{0,XY}$  and  $Q_{1,00}$ ,  $\{i \mid BT1231_{Q_0,XY} \leq BT1231(i) \leq BT1231_{Q_{1,00}}\}$ . In what follows in this subsection we use an "integrated" or "cumulative" quantile wherein Q0.90 now means all scenes between Q0.90 and Q1.00 (maximum observed BT) are considered. XY is the radiances averaged over all the observations *i* which are in the set  $\Psi_{0,XY}$ , namely

$$r_{Q0,XY}(\nu) = \frac{1}{N_{0,XY}} \sum_{i \in \Psi_{0,XY}} r_i(\nu)$$
(1)



**Figure 1.** Zonally averaged BT1231 normalized histograms (probability distribution functions) as a function of latitude and temperature bin, for an 2012/08/27 - 2012/09/11 timespan (colorbar) and quantiles (curves). The thick black curve is the *Q*0.90 quantile (and above) used in this paper, and is very close to the maximum *Q*1.00 quantile.

where  $r_i(\nu)$  are the  $N_{0,XY}$  individual observations in set  $\Psi_{0,XY}$ . In this section we only use the  $\nu = 1231 \text{ cm}^{-1}$  channel, but in later sections we easily form averages for all 2645 channels, at any 16 day time step for any tile.

To further investigate if the scenes chosen using this definition can be considered 367 We tested different quantile sets  $\Psi_{0,XY}$  to see which one can reliably be considered to 368 provide a nominally "cloud free", we compare to global dataset, and chose the Q0.90 average 369 (ie defined as averaged over the  $\Psi_{0.90}$  set) as the one to use for the rest of this paper, 370 unless explicitly stated otherwise. The tests primarily involved comparisons to scenes 371 produced by the uniform/clear sky filter described in Strow and DeSouza-Machado (2020) 372 for the same August 27, 2012 - September 11, 2012 sixteen day timespan. This latter fil-373 ter selects clear scenes by both testing for uniformity (to within 0.5 K) across a  $3 \times 3$ 374 grouping of AIRS scenes and also using a criteria that the observed window channel ob-375 servations should be within  $\pm 4$  K of clear-sky simulations using thermodynamic param-376 eters supplied by reanlysis reanalysis models. The results are shown in the left hand plot 377 of Figure 2, plotted on a  $1^{\circ} \times 1^{\circ}$  grid. We note in this plot the uniform/clear scenes that 378 are plotted are limited to those over ocean, and for solar zenith less than 90  $^{\circ}$  (daytime), 379 which automatically filtered out many of the views over the (wintertime) Southern Po-380 lar region. Immediately apparent are the gaps produced by the uniform/clear filter e.g.381 in the Tropical West Pacific or off the western coasts of continents where there are clouds. 382 The gaps can be changed by e.g. changing the 4K threshold to allow more or fewer scenes 383 through the filter. 384

The center plot shows the scenes selected by the integrated for all tiles, the daytime scenes selected for the Q0.90 filter average for the same time period, on the same  $1^{\circ} \times$  $1^{\circ}$  grid. Compared to the left hand plot, the spatial coverage is almost complete, as the Q0.90 filter average always has the hottest 10% of the observations, the spatial coverage is almost complete : gaps are only visible. At this 1° resolution, used for comparison with the uniform/clear grid filter described in the previous paragraph, gaps are seen in regions



Figure 2. Clear scenes for the same 2012/08/27 - 2012/09/11 timespan selected by (left) an uniform/clear sky filter and (center) the Q0.90 integrated filter average described in this paper. The right hand plot shows the mean (over ocean) observed BT1231 as a function of latitude, for the two selections; the difference is about 0 K  $\pm$  1 K in most region except in the southern mid-latitudes where the integrated-Q0.90 filter average produced scenes that were about 1 K cooler on average.

where there are for example mountains, or in the desert regionswhere other areas are even warmer. We note that increasing the quantile threshold to 0.95 or 0.97 did not introduce the gaps seen in the left hand (uniform/clear) mapfor example the local topography means observations over mountains would be colder than the surrounding coastal or plain regions. This is not a concern since zooming back out to the coarser  $3^{\circ} \times 5^{\circ}$  tile resolution, will include Q0.90 data for the quantile and trending analysis.

To compare the mean observations we filter away remove the over-land and over-397 polar region data from the center plot. The right hand plot shows the mean observed 398 BT1231 from the  $1^{\circ} \times 1^{\circ}$  grid from the uniform/clear sky filter as a function of latitude, 399 compared to the  $1^{\circ} \times 1^{\circ}$  grid from the integrated Q0.90 scenes. The difference between 400 the uniform/clear versus integrated Q0.90 filter average is within about 0.25 K  $\pm$  1 K 401 across the southern tropics to the northern midlatitudes, though the bias rises to about 402 1 K by about  $-50^{\circ}$ S. We consider this an acceptable difference, as we could tune the thresh-403 olds for the uniform/clear filter to e.q. change the areal coverage and/or number of clear 404 scenes and hence comparisons to the Q0.90 scenes. 405

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We ran these tests for three. The results presented in this section have been checked 407 for robustness, using other 16 day intervals in 2012, spanning the four seasons. The overall 408 global bias and standard deviation for the 1231 channel between the co-located  $1 \times 1$  grids 409 by the uniform/clear filter and by the Q0.90 filter stayed fairly uniform, typically about 410  $0.25 \text{ K} \pm 1 \text{ K}$ . From the information presented in this section, we We conclude that for 411 any 16 day timestep the integrated radiances used in the Q0.90 filter average (a) pro-412 duces almost complete spatial coverage of the Earth, (b) selects scenes whose average 413 BT1231 is very close to the average BT1231 from scenes selected using an uniform/clear 414 filter (c) trends from that quantile typically differ by less than  $\pm 0.002 \text{ K} / \text{yryr}^{-1}$  from 415 the other quantiles and (d) this selection produces spectral trends which compare well 416 against those obtained from the quality assured binned AIRS CCR data record Manning 417 (2022). Together these imply the integrated Q0.90 average is an acceptable proxy for "clear 418 scenes". For the remainder of the paper we drop the word "integrated" and therefore con-419 sider Q0.90 as consisting of nominally clear observations whose BT1231 lies between the 420 90th quantile and hottest observation. Our retrievals using this  $\mathcal{Q}0.90 \rightarrow \mathcal{Q}1.00$  averaged 421 dataset (shortened to Q0.90) is referred to as AIRS RTin what follows. 422

# 3.2 Observed trends from the Q0.90 Quantiles

Having selected the Q0.90 observations, for each tile the average radiance per 16 day interval is computed. With two sixteen day periods not available (Aqua platform or AIRS shutdowns during *e.g.* solar flare events) this gives a total of 457 time steps over 20 years. Anomalies are formed from this time series, and then de-seasonalized to give the spectral radiance trends and error estimates Strow and DeSouza-Machado (2020) using Matlab *robustfit*:

$$r_{\text{observations}}^{16 \text{ days}}(t) \sim r_{\text{fit}}(t) = r_o + a_1 t + \sum_{i=1}^4 c_i sin(n2\pi t + \phi_i)$$
 (2)

with  $a_1$  and its associated uncertainty, both converted to brightness temperature (BT), being the trends in K /yryr<sup>-1</sup>. Using sub-harmonics in the fit did not produce any noticeable change in the AIRS RT retrievals (described below).

The left panel of Figure 3 shows the descending orbit (nightime) 20 year (Septem-433 ber 2002- August 2022) global averaged spectral observations for the five quantiles men-434 tioned above. We note the spectra in most of the plots in this section are weighted by 435 the cosine(latitude) of the tiles, unless otherwise stated. In addition we only show the 436  $640-1640 \text{ cm}^{-1}$  region, and ignore the shortwave 2050-2750 cm<sup>-1</sup> region since the AIRS 437 SW channels are drifting relative to the LW Strow and DeSouza-Machado (2020). Spec-438 tral averages constructed from Figure 1 would have this same behavior, namely that in 439 the window region the mean spectrum of <u>data populating</u> the warmer quantiles integrated 440 out to Q1.00 (Q0.80, Q0.90, Q0.95, Q0.97) as defined in Equation 1 are on the order of 441 a Kelvin apart, and have about half/quarter that difference in the optically thicker re-442 gions dominated by  $H_2O_{and}/or_{and$ 443

The right hand panel of Figure 3 shows (top) the trends and (bottom) the  $2\sigma$  trend 444 uncertainties for these quantiles, in  $\frac{\text{Kelvin/year}}{\text{Vear}} \text{K} \text{yr}^{-1}$ . We emphasize that the top right 445 panel shows that the spectral trends for the quantiles lie almost on top of each other; 446 the difference between the Q0.50 and other trends is at most about +0.003 K  $/yryr^{-1}$ 447 (out of a 0.02 K  $/yryr^{-1}$  signal) in the window region (and about +0.0045 K  $/yryr^{-1}$ 448 in the troposphere temperature sounding channels), or less than 10%. Similarly the largest 449 trend uncertainty in the bottom panel is for  $Q_{0.50}$ . This implies that clouds effects in 450 the infrared do produce the largest variability (blue curve) but on average for the infrared 451 are not changing much, so the  $+0.022 \text{ K} / \text{year} \text{yr}^{-1}$  window region trends are dominated 452 by surface temperatures changes and to a lesser extent by water vapor changes. 453

The TOA radiances in the 15 um (700-800 cm<sup>-1</sup>) region is dominated by the are 454 impacted by two effects (a) the increased optical depths due to increasing atmospheric 455 CO<sub>2</sub> increases; the effects of increasing are to make the atmosphere emit at leads to atmospheric 456 emission from higher altitudes/lower temperatures, leading to an almost resulting in almost 457 a -0.06 K/year signal for the troposphere; hidden in there are the, and (b) the atmospheric 458 temperature increases (again about +0.02 K  $\frac{1}{(year)}$ ; also yr<sup>-1</sup>). Also of interest is the 459 trends in the stratosphere (650-700  $\rm cm^{-1}$ ) changes which consists of a stratospheric cool-460 ing signal (negative) and emission higher up due to increased CO<sub>2</sub>; combining to give 461 a net zero effect over 20 years, also seen in Raghuraman et al. (2023). The  $H_2O$  signal 462 is evident in the 1400-1625  $\rm cm^{-1}$  region, and is negative; in other words, increasing tem-463 peratures have led to increased atmospheric amounts of  $H_2O$ , and the water vapor feed-464 back has reduced the amount of outgoing flux in that region. By extension, this also hap-465 pens in the Far Infrared regions affected by water vapor; current sounders do not make 466 direct measurement in the  $10-600 \text{ cm}^{-1}$  region so at present this can only be inferred; how-467 ever in the near future it is anticipated the Far Infrared Outgoing Radiation Understand-468 ing and Monitoring (FORUM) mission Palchetti et al. (2020) will provide data to fill in 469 this important gap in the future. 470



Figure 3. 20 year trends from different observation quantiles. The left hand panel shows the mean globally averaged BT trends observations from 20 years of AIRS data, for quantiles 50Q0.50,800.80,900.90,950.95,97–0.97 as described in the text. The right hand panel shows (top) the globally averaged trends from for those years different quantiles and (bottom) the spectral uncertainty in the trends. The nightime (descending) trends are shown in these plots.

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### 3.3 Observed trend changes over 20 years

The left panel of Figure ?? shows the mean descending (nightime) orbit Q0.90 brightness 473 temperature spectrum, for four time periods, all commencing on September 1, 2002 - the 474 periods are for 5,10,15,20 years of data and end on August 31, 2007, 2012, 2017, 2022 respectively. 475 As expected the mean cosine averaged observed BT is slightly over 284 K through most 476 of the longwave window region. The right hand panel of the same Figure ?? shows the 477 trends for the four time periods in the top, while the bottom shows the uncertainties. 478 Averaging over the inter-annual variability affects the trends, with the shortest/longest 479 time periods (5/20 years) having the largest/smallest spectral uncertainty as one would 480 expect as inter-annual variability slowly becomes less important in the trends. 481

Changes in AIRS observations over time spans of 05,10,15,20 years all beginning
on September 1, 2002. The left hand panel shows the mean globally averaged 90th quantile
BT spectra for those time periods. The right hand panel shows (top) the trends from
those years and (bottom) the spectral uncertainty in the trends. The nightime (descending)
trends are shown in these plots.

# 487 4 Spectral closure : comparisons between observed and simulated spec-488 tral trends

Previous work Strow and DeSouza-Machado (2020) has demonstrated that the ra-489 diances from AIRS are climate quality, if one restricts the channel set to the  $\sim 450$  chan-490 nel set that is largely immune to nonphysical drifts Strow et al. (2021). In this section 491 we describe a way to test the quality of the monthly thermodynamic output from reanal-492 ysis and/or L3 products which are all in geophysical space, against the AIRS L1C ob-493 servational data which is in radiance space. This is accomplished by geolocating the monthly 494 (ERA5) surface temperature, air temperature, water vapor and ozone fields to tile cen-495 ters as described in Section 2.2, which are then input and run through the SARTA fast model Strow, Hannon, DeSouza-Machado, et al. (2003), for the entire 20 years. Spec-497 tral radiance trends were then computed from these time series of (clear sky) spectral 498 radiances. The conversion of L3 retrieval and NWP reanalysis trends to a radiance time 499



Figure 4. 20 year zonally averaged spectral brightness temperature trends (in K  $/yearyr^{-1}$ ) for (left) AIRS Q0.90 observations and (right) clear sky simulations using ERA5 monthly model fields. The center panel shows the AIRS Q0.90 spectral uncertainties. The ERA5 simulations included linear trends of CO<sub>2</sub>, CH<sub>4</sub> and N<sub>2</sub>O, while the O<sub>3</sub> trends in ERA5 are from the reanalysis itself.

series, provides a rigorous check of their accuracy against the observed AIRS L1C ra diance trends which are validated to be highly accurate.

A good reviewer might ask about the noise introduced by secant angle varying in the 16 day period. Check

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The simulations included realistic column linearly-increasing-with time mixing ratios for  $CO_2$ ,  $CH_4$  and  $N_2O$  for the ERA5 spectra, as well as land or ocean surface emissivity co-located to tile centers together with view angles of about 22°. From these the ERA5 spectral trends were derived similarly to what was described above for the AIRS observation spectral trends.

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Figure 4 shows the descending (night) zonally averaged results in K  $\frac{1}{2}$  year yr<sup>-1</sup>, al-510 lowing us to compare the Q0.90 nominally clear AIRS observed spectral trends, to those 511 simulated using monthly ERA5 fields (without clouds). The center panel shows the spec-512 tral trend uncertainties from the observations, also in K  $\frac{1}{\sqrt{2}}$  we have the next section 513 we derive geophysical trends from these (AIRS observed) spectral trends, and the sim-514 ilarities/ differences in geophysical trends can be partially understood from the similar-515 ities/differences in the spectral trends. For example, the  $H_2O$  sounding region (1350-1600) 516  $cm^{-1}$ ) shows roughly similar (positive) trends in the tropics and mid-latitudes; there are 517 some slight differences in the high altitude channels (1450-1550  $\rm cm^{-1}$  region). The fol-518 lowing sections shows that there are subtle differences in these trends, which manifest 519 as differences in tropospheric water vapor trends. Observations and simulations both have 520 positive dBT/dt in the 800-960,1150-1250 cm<sup>-1</sup> region, indicating surface warming; how-521 ever the ERA5 simulation show more warming in the southern polar regions than do the 522 AIRS observations. In particular note the mean warming in the tropics is less than that 523 in the mid-latitudes, and the polar regions show the largest overall change in brightness 524 temperature in the window region. Large differences are seen in the 10 um  $(1000 \text{ cm}^{-1})$ 525  $O_3$  sounding region, which are not surprising since ozone assimilation is not a primary 526 goal of ECMWF assimilation; here we do not address these as we focus on the changes 527 to the moist thermodynamic state. The window region trends computed using the ERA5 528 model are more positive in the Southern Polar region. Conversely the 640-700  $\rm cm^{-1}$  spec-529 tral region is positive, especially in the tropics; however the observations show a net cool-530 ing trend away from the tropics, compared to the ERA simulations. This demonstrates 531 the importance of the model  $\rightarrow$  spectral trend comparisons, given the accuracy of the 532 AIRS observations. 533

The paper by (Raghuraman et al., 2023) shows similar figures, but in terms of spec-534 tral OLR trends encompassing the 0-2000  $\rm cm^{-1}$  range, while (Huang et al., 2023) shows 535 similar plots for a slightly smaller time period (2002-2020) and using nadir L1B radiance 536 dataset which has no or minimal frequency corrections compared to the L1C set we use 537 in this paper. (Huang et al., 2023; Raghuraman et al., 2023) and our work all show, ei-538 ther in radiance or OLR space, (a) the increased observed radiance in the window chan-539 nels, due to surface temperature increases (b) the  $\simeq$  -0.06 K /yryr<sup>-1</sup> decrease in BT in 540 the 700-750  $\rm cm^{-1}$  troposphere sounding region, which is due to the CO<sub>2</sub> amounts increas-541 ing; we also see differences in the signs of the BT changes in the 650-700  $\rm cm^{-1}$  strato-542 spheric CO<sub>2</sub> and temperature channels for some latitudes between AIRS RT observa-543 tions and ERA5 simulations (c) increases in the 1350-1640  $\rm cm^{-1}$  water vapor sounding 544 region seen in Figures 3 and 5, and (d) the 1280-1340  $\rm cm^{-1}$  decreases are due to  $\rm CH_4$ 545 increases. 546

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# 4.1 Sample spectral closure comparisons using other monthly products

Here we follow the earlier work of Huang et al. (2023) and convert the ERA5 monthly 548 model fields to spectral radiances, after which we compute spectral trends for compar-549 ison to AIRS observations. Spectral closure calculations for the entire 20 year timeseries 550 were also generated for the monthly MERRA2 model fields, as well as the monthly AIRS 551 v7 L3 and CLIMCAPS L3 retrieved data products. Again only the monthly thermody-552 namics and surface temperature fields for all  $72 \times 64$  tiles were used in these SARTA runs, 553 with GHG changes added in for each timestep as described above. Spectral trends were 554 then computed using Equation 2. 555

We chose just one limited example here to illustrate the power of this approach for 556 diagnosing which dataset is more accurate, given that the AIRS spectral trend accuracy 557 is already established. Water vapor is highly variable in space and time, meaning wa-558 ter vapor retrievals using hyperspectral sounders radiances differ most from NWP fore-559 casts, in particular because of the typical  $\pm$  90 minute difference between observation 560 and forecast, and is where these sounders typically provide the most information. Fig-561 ure 5 show the globally averaged brightness temperature trends (in K /yearyr<sup>-1</sup>) in the 562  $1350 - 1650 \text{ cm}^{-1}$  water vapor sounding region. The blue curve shows the trends from 563 the AIRS observations used in this paper, while spectral trends constructed from the AIRS 564 L3/ CLIMCAPS L3 retrievals are in red/yellow and the ERA5 model fields are in pur-565 ple. The AIRS observations and ERA5 constructed spectral trends are positive in this 566 region, while the AIRS L3 and CLIMCAPS L3 trends are obviously different, being neg-567 ative in this water vapor sounding region. The subtle differences in these spectral trends 568 arise from differences in the geophysical trends between observations and the models them-569 selves, and will be addressed in the following sections, where the retrieved and model sur-570 face temperature, and atmospheric temperature and water vapor geophysical trends will 571 be compared and discussed. 572

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# 574 5 Testing the variability of representative points from NWP modelsreanalysis

Each sixteen day  $3^{\circ} \times 5^{\circ}$  tile contains ~ 12000 observations, which means for each tile about 600 daytime and 600 nightime observations are averaged to produce the Q0.90 dataset per timestep. Conversely there are typically only ~ 240 monthly ERA5 0.25° points per  $3^{\circ} \times 5^{\circ}$  tile; for 1° resolution AIRS L3 and CLIMCAPS L3 there are even fewer (15) points per tile. This low number of points means we chose a simple solution of using the grid cell closest to the center of each  $3^{\circ} \times 5^{\circ}$  tile for building the NWP and L3 geophysical time series. This choice is validated below using the following test to see for exam-



Figure 5. Globally averaged spectral trends in the water vapor sounding region : AIRS L1C observations (blue) compared to spectral closure from the standard monthly AIRS L3 retrievals (red) and CLIMCAPS L3 (yellow) and from monthly ERA5 simulations (yellow). The reconstructed AIRS\_RT trends very closely match the AIRS L1C observations and are not shown here.

ple how surface temperature trends would be impacted as we changed the representative point for the ERA5 model fields.

For the descending overpass we built complete sets of approximately 240 ERA5 points 584 per tile per month; at 0.25° resolution one of these is almost certainly at the tile center. 585 From these monthly sets, we could either directly read the tile center temperature (our 586 default), or compute the average surface temperature per tile, or compute the average 587 of the hottest 10% surface temperatures per tile. This was done for all 20 years (240 monthly 588 timesteps) after which the three timeseries were trended. Over ocean the differences be-589 tween all three datasets as was typically  $-0.001 \pm 0.005$  K /yearyr<sup>-1</sup>, while over land the 590 differences were larger at about  $0.001 \pm 0.01$  K /yearyr<sup>-1</sup>. This is to be compared to 591 mean trends of about 0.014  $\pm$  0.02 K  $/yryr^{-1}$  over ocean and 0.025  $\pm$  0.04 K /yrover592 land. In other words yr<sup>-1</sup> over land : the spread of the ocean and land ERA5 surface 593 temperature trends for the three methods, was about four times larger than the spread 594 of the differences between the three methods. In what follows is much smaller than the 595 mean trends. Given that there were far fewer re-analysis points in a grid box than tiled 596 Q0.90 observations, coupled with the fact that choosing the 10% warmest profiles would 597 provide an even smaller sample, we chose to use the tile center was thus chosen as to be 598 the representative point to co-locate the model fields<del>, when comparing against the tiled</del> 599 observations. 600

# 6 Geophysical Trend Retrieval outline

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# 6.1 Setting up the Retrieval Problem

The observed spectral brightness temperature for a tile at any time t can be modeled as

$$BT(\nu, t) = f(X(t), \epsilon(\nu, t), \theta(t)) + \text{NeDT}(\nu)$$
(3)

where the state vector X(t) has the following five geophysical state parameters : (1) surface temperature (ST), (2) atmospheric temperature profile T(z), (3) water vapor pro-

file WV(z), (4) ozone profile O3(z) (5) greenhouse gas forcings (GHG) due to CO<sub>2</sub>, CH<sub>4</sub> 607 and N<sub>2</sub>O changing as a function of time t and  $f(X(t), \epsilon, \theta, \nu)$  is the clear sky radiative 608 transfer equation for channel center frequency  $\nu$ . The spectral noise NeDT( $\nu$ ) for a typ-609 ical tropical "clear scene" is about 0.1 K in window region, increasing to about 1 K in 610 the 15  $\mu m$  temperature sounding channels and about 0.2 K in the 6.7  $\mu m$  water vapor 611 sounding region, but the noise will vary as a function of scene temperature. We parameterize 612 parametrize the GHGs using single numbers (such as ppm(t) for the CO<sub>2</sub> column), and 613 include the AIRS orbit and viewing angle geometry  $\theta$  and the surface emissivity  $\epsilon(\nu)$ , 614 while we omit forward model and spectroscopy errors. We ignore cloud scattering as well 615 as the spatial variation of the state parameters, emissivity and scan angle geometry within 616 a tile. Linearizing the above equation about the time averaged profile, the relationship 617 between the observed spectral trends and desired thermodynamic trends is given by 618

$$\frac{d\overline{BT(\nu)}}{dt} = \frac{\partial f}{\partial \overline{X}} \frac{d}{dt} \overline{X(t)} = K(\nu) \frac{d}{dt} \overline{X(t)} + \underbrace{K_{\text{emissivity}}(\nu)}_{dt} \frac{d}{dt} \overline{\epsilon(t)} \xrightarrow{0} K(\nu) \frac{d}{dt} \overline{X(t)}$$
(4)

where the matrix  $K(\nu)$  is the thermodynamic jacobian (surface temperature, air temperature and trace gases) and we ignore any orbit drifts (changes to  $\theta$ ), instrument changes (changes to  $NeDT(\nu)$ ) and surface emissivity ( $\epsilon(\nu)$ ); the last assumption is investigated in a later section. The overbars on parameters X denotes this is a time average (linear trend) that we are working with, and we have converted from radiances in Equation 2 to brightness temperatures in Equations 3 and 4.

# 6.2 Jacobian calculations

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For a typical clear sky tropical sky atmosphere, the  $800 - 1200 \text{ cm}^{-1}$  window re-626 gion has surface temperature (SKT) jacobians which are about +0.5 to +0.75 K per de-627 gree SKT change and -0.75 to -0.25 K per 10% change in column water vapor. The spec-628 tral variability in these window region jacobians is primarily due to reducing water con-629 tinuum absorption as you move from the 800 cm<sup>-1</sup> end to the 1200 cm<sup>-1</sup>; consequently 630 the surface temperature jacobians becomes closer to unity and the column water jaco-631 bians become closer to zero as water vapor amount decreases (drier atmospheres in the 632 mid-latitudes and polar regions). The hyperspectral channels used in this work help sep-633 arate out these two competing changes, which we validate against other datasets in this 634 study. As seen in Figure 4 typical magnitudes of the spectral trends on the left hand side 635 of Equation 4 are less than about 0.1 K per year. Equation 4 is in the usual inversion 636 form  $\delta y = K \delta x$ , and the Optimal Estimation (Rodgers, 2000) solution used to solve 637 the anomaly time series in Strow et al. (2021) is also used here. The noise term used for 638 the trend retrieval  $NeDT(\nu)$  is not the instrument noise since each 16 day point in our 639 time series is averaged over hundreds of observations as earlier described; instead the un-640 certainty is that due to inter-annual variability in the linear trends obtained from the 641 trend fitting in Equation 2. Examples of typical noise values are shown in the bottom 642 right hand panels panel of Figures 3and ??. 643

ERA5 monthly model fields at tile centers, together with time varying concentra-644 tions of GHG such as  $CO_2$ , were averaged over 20 years so jacobians could be computed. 645 The GHG concentrations were a latitude dependent increase of about  $\sim 2.2 \text{ ppmv/year}$ 646  $ppm yr^{-1}$  for CO<sub>2</sub> derived from the CarbonTracker Peters et al. (2007) (CarbonTracker 647 CT-NRT.v2023-4, http://carbontracker.noaa.gov) model data at 500 mb. Our pseudo-648 monochromatic line by line code kCARTA De Souza-Machado et al. (2018, 2020) was used 649 with these averaged profiles to produce accurate analytic jacobians. The HITRAN 2020 650 line parameter database Gordon and Rothman (2022), together with MT-CKD 3.2 and 651  $CO_2, CH_4$  line mixing from the LBLRTM suite of models Clough et al. (2005) were used 652 in the kCARTA optical depth database De Souza-Machado et al. (2018). A 12 month ge-653 ographical land-varying spectral emissivity database spanning one year from Zhou et al. 654 (2011) was used, while ocean emissivity came from Masuda et al. (1988). The atmospheric 655

temperature, water vapor and ozone profile jacobians, and the surface temperature and column jacobians for the GHG gases such as  $CO_2$  and  $CH_4$  and  $N_2O$ , were then convolved using the best estimate AIRS Spectral Response Functions Strow, Hannon, Weiler, et al. (2003).

Tests done for this paper, together with the results in Strow et al. (2021), established that jacobians derived from MERRA2 versus ERA5 produced no significant differences in the context of retrieved trends or anomalies done for this paper, as the uncertainty in linear trends due to inter-annual variability dominates over any uncertainty (or differences between) model fields.

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# 6.3 Optimal Estimation Retrieval : State vector, covariance matrices and *a-priori*

Using monthly ERA5 model fields averaged over 20 years, for each of the  $64 \times 72$ 667 tiles we computed analytic jacobians for the following (vector) atmospheric thermody-668 namic variables [fractional water vapor, fractional ozone and temperature] together with 669 (scalar) surface temperature, where we retrieved fractional gas concentration trends dfracX/dt =670  $1/X_{avg}(z)dX_{avg}(z)/dt$  to keep all values in the state vector at about the same magni-671 tude. A single iteration Optimal Estimation retrieval Rodgers (2000) is used to simul-672 taneously solve for the geophysical parameter trends. As in (Strow & DeSouza-Machado, 673 2020) the geophysical covariance uncertainty matrices are a combination of Tikonov and 674 covariance regularization. The uncertainties for the covariance matrices were typically 675 [0.1, 0.25, 0.45] K /yryr<sup>-1</sup> for the surface/tropospheric/stratospheric temperature trends, and [0.04/0.02] /year yr<sup>-1</sup> for the fractional tropospheric/stratospheric water vapor trends. 676 Tikonov L1 regularization (Rodgers, 2000) also included, with the scalar factor multi-678 plying this regularization corresponding to about 1/10 the covariance uncertainties. The 679 spectral uncertainties used in the retrievals come from the above mentioned trend un-680 certainties. For completeness we note that a sequential retrieval (see for example Smith 681 and Barnet (2020)) produces very similar geophysical trends. 682

Here we emphasize four points about our geophysical trend retrievals, which sets 683 us apart from trends derived from other datasets. Firstly the *a-priori* trend state vec-684 tor is zero  $\frac{dST}{dt} = \frac{dT(z)}{dt} = \frac{dQ(z)}{dt} = 0$  for all geophysical parameters, except 685 for water vapor where we enforced constant (or slightly increasing) relative humidity as 686 described below. This ensures traceability of our retrieval is straightforward especially 687 wherever the AIRS instrument has sensitivity. For example the 300 - 800 mb water va-688 por trend retrievals will be based on the data only, thereby insulating us from any pos-689 sible *a-priori* information from *e.q.* climatology or NWP models, unlike the operational 690 AIRS V7 or CLIMCAPS retrievals which use first guesses based on neural net and MERRA2 691 respectively. 692

Secondly as seen in Figures 4 and 5, in the 15  $\mu m$  region there is a large spectral 693 overlap signal (-0.06 K  $/yryr^{-1}$ ) from the increasing CO<sub>2</sub>, which is much larger than the 694 expected atmospheric temperature trend (0.01 K /yr). The 20 year dataset contains inter-annual 695 variability whose noisy time series and correlations with for example temperature changes, 696 which  $yr^{-1}$ ). These correlations makes it difficult to also retrieve these well mixed GHG. 697 Instead of attempting to solve for both GHG concentration changes and for temperature 698 <del>changes, we spectrally removed</del> jointly retrieve both temperatures changes and changes 699 in well mixed GHGs such as CO<sub>2</sub>. We chose to focus on retrieving temperature changes 700 only, by spectrally removing the effects of changing  $\frac{\text{ed}CO_2}{\text{CO}_2}$ , CH<sub>4</sub> and N<sub>2</sub>O GHG con-701 centrations, This was done by using the GHG trends estimated from NOAA ESRL Car-702 bonTracker data multiplied by the appropriate GHG gas column jacobian  $(CO_2, N_2O)$  and 703  $CH_4$  and CFC11, CFC12) computed as described above using the averaged over 20 years 704 ERA5 monthly profile for each tile. 705

Thirdly instead of using all 100 layers described in the AIRS forward model Strow, Hannon, DeSouza-Machado, et al. (2003), we combine pairs of layers for a 50 atmospheric layer retrieval, as the AIRS radiances contain far fewer than 100 pieces of information (see *e.g.* Maddy and Barnet (2008); De Souza-Machado et al. (2018)).

Fourthly, modern hyperspectral infrared sounders have highest sensitivity to tem-710 perature and water vapor in the mid-tropopause; see for example the averaging kernels 711 in Irion et al. (2018). Using a zero fractional WV trends *a-priori* at all levels, it was fairly 712 straightforward to obtain fractional WV(z) trends close to those from the NWP model 713 datasets in the <del>3000-850</del> 300-850 mb region. In order to improve our results in the low-714 est layers, we enforced a constant relative humidity approximation, which is a well-known, 715 expected behavior under global climate change Soden and Held (2006); Sherwood et al. 716 (2010). This was done by using the ignoring the contribution due to water vapor changes 717 in the observed BT1231 trend, and using it as an approximation for air temperature trend 718 over ocean; this allows us to compute an estimate of how the water vapor would need 719 to change 720

$$RH(T) = \frac{e}{e_{sat}(T)} \implies \delta(RH) = \frac{1}{e_{sat}(T)} \delta e - \frac{e}{e_{sat}^2(T)} \delta e_{sat}(T) = \frac{1}{e_{sat}(T)} \delta e - \frac{e}{e_{sat}(T)} \frac{L_v}{R_v} \frac{1}{T^2} \delta T$$
(5)

where  $e, e_{sat}(T)$  are the vapor pressures and we used  $e_{sat}(T) = e_{s0}e^{\frac{L_v}{R_v}\left(\frac{1}{T_o} - \frac{1}{T}\right)}$  (where 721  $L_v, R_v$  are latent heat of vaporization and gas constant respectively) to go from the ex-722 pression in the center to the expression on the right. If we expect the change in RH to 723 be zero then  $\frac{\delta e}{e} = \frac{L_v}{R_v} \frac{\delta T}{T^2}$ , where we can use  $\delta T/\delta t \sim d/dt BT 1231$ . to approximate the 724 a-priori fractional vapor pressure rates (or a-priori fractional water vapor rates) between 725 surface and 850 mb, smoothly tailing to 0 in the upper atmosphere. Subsection 7.2 has 726 a similar discussion on a proposed method to alleviate the lack of sensitivity to upper 727 atmosphere water vapor. Our default results in this paper are from using the MLS a-728 priori, unless otherwise stated. 729

730 731

# 6.4 Testing on Synthetic Spectrasynthetic trend spectra made from ERA5 Reanalysis monthly fields

We tested the retrieval code by using it on the simulated nighttime only ERA5 spec-732 tral trends, and compared to geophysical trends computed directly from the ERA5 re-733 analysis model. Spot checks of the spatial correlations of ERA5 fractional water vapor 734 and temperature trends versus the trends retrieved from synthetic spectra/our retrieval 735 algorithm, peaked at 500 mb with correlations of about 0.9, compared to 800 mb cor-736 relations of 0.80 and 0.55 for temperature and fractional water vapor trends respectively 737 and 200 mb correlations of 0.89 and 0.69 for dT/dt, dWVfrac/dt. This is to be expected 738 since a computation of the water vapor averaging kernels for infrared instruments for ar-739 bitrary atmospheric profiles typically shows they peak in the 300 mb - 850 mb range and 740 decrease rapidly away from those regions; conversely the temperature averaging kernels 741 stay relatively uniform through the free troposphere and above, though they also decrease 742 close to the surface; see for example Irion et al. (2018); Smith and Barnet (2020); Wu 743 et al. (2023). 744

Figure 6 shows a sample set of results using nightime ERA5 model output converted to spectral trends as described above. The top panels (A) are always the atmospheric trends derived direct computed directly from the monthly ERA5 model fields, while the bottom panels (B) are the atmospheric trends derived retrieved from the converted ERA5 spectral brightness temperature trends. The left most panel is the atmospheric temperature trend comparison (both in K /yryr<sup>-1</sup>) while the rightmost panel is the fractional atmospheric water vapor trend comparison (in /yryr<sup>-1</sup>).



Figure 6. Comparing geophysical trends derived directly from ERA5 monthly nighttime fields (top) vs from the OEM retrieval applied to the spectral trends (bottom). Left panel is dT/dt (in  $K \neq yearyr^{-1}$ ) while rightmost panel is d(fracWV)/dt (colorbar in  $\neq yr^{-1}$ ).

753	It is evident from the figure that the tropospheric trends in the tropical and mid-
754	latitude regions are quite similar, and there are differences in the polar regions and strato-
755	spheric regions where the AIRS instrument has reduced sensitivity. The atmospheric and
756	surface trends are shown in Table 1, divided into "all" (which is the entire $\pm$ 90 latitude
757	range and 0-1000 mb vertical range) and "T/M" which is the tropical/midlatitude region,
758	which is further reduced to 050-900 mb for air temperature and 300-800 mb for water
759	vapor. "ERA5 direct" are trends computed directly from the geophysical fields, while "ERA5
760	spectral" are retrieved from the spectral trends.

	$\begin{array}{c c} & dTz/dt \\ & K \frac{/yryr^{-1}}{A} \\ \hline & A \\ \hline & GND-TOA-SFC-TOA \end{array}$	$\begin{array}{c} \mathrm{dTz/dt} \\ \mathrm{K} \frac{/\mathrm{yryr}^{-1}}{\mathrm{T/M}} \\ \mathrm{050\text{-}900 \ mb} \end{array}$	$\begin{vmatrix} dSKT/dt \\ K {/yryr^{-1}} \\ A \end{vmatrix}$	$\begin{array}{c} \mathrm{dSKT}/\mathrm{dt} \\ \mathrm{K} \frac{/\mathrm{yryr}^{-1}}{\mathrm{T/M}} \\ \mathrm{T/M} \end{array}$	$ \begin{array}{c c} \mathrm{dfracWV}/\mathrm{dt} & \\ \mathrm{K} \frac{/\mathrm{yryr}^{-1}}{\mathrm{A}} \\ \mathrm{GND}\text{-}\mathrm{TOA} \end{array} $	dfracWV/ K <del>/yryr</del> T/M 300-800 n
ERA5 direct ERA5 spectral	$\begin{array}{c} 0.010 \pm 0.038 \\ 0.004 \pm 0.033 \end{array}$	$\begin{array}{c} 0.029 \pm 0.013 \\ 0.027 \pm 0.012 \end{array}$	$\begin{vmatrix} 0.020 \pm 0.035 \\ 0.019 \pm 0.033 \end{vmatrix}$	$\begin{array}{c} 0.018 \pm 0.032 \\ 0.016 \pm 0.029 \end{array}$	$ \begin{vmatrix} 0.003 \pm 0.002 \\ 0.001 \pm 0.001 \end{vmatrix} $	$0.002 \pm 0.0$ $0.002 \pm 0.0$

**Table 1.** Cosine weighted air temperature, skin temperature, fractional water vapor trends, together with uncertainties/. The "ERA5 direct" are directly from the ERA5 geophysical trends, while "ERA5 spectral" are trends retrieved from the converted ERA5 spectral trends.

# 6.5 Surface emissivity changes

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Equation 3 explicitly includes the surface emissivity in the equation of radiative
 transfer; however Equation 4 assumes this is unchanging. Here we rewrite Equation 4
 as

$$\frac{d\overline{BT(\nu)}}{dt} - K_{emissivity}(\nu)\frac{d}{dt}\overline{\epsilon(t)} \to \frac{d\overline{BT'(\nu)}}{dt} = K(\nu)\frac{d}{dt}\overline{X(t)}$$
(6)

 $_{765}$ First we consider ocean emissivity changesOcean emissivity has a dependence on $_{766}$ windspeed Masuda et al. (1988). Lin and Oey (2020) and other literature suggest wind $_{767}$ speed increases of +2.5 cm  $\frac{/s/year s^{-1} yr^{-1}}{1}$  have occured between 1993-2015 in the trop- $_{768}$ ical Pacific, and smaller (or close to zero) values elsewhere. The monthly ERA5 u10,v10

10 m speeds for the 20 year time period in this paper also showed the maximum abso-769 lute trend was 0.09 m/s/year (over the Southern Ocean) while the global ocean mean 770 and standard deviation were  $0.006 \pm 0.022$  m  $\frac{\text{/s/years}^{-1} \text{ yr}^{-1}}{\text{/s}^{-1} \text{ yr}^{-1}}$ ; The emissivity changes 771 over ocean using a 0.025 m  $\frac{1}{1000}$  wind speed change are on average on the order of 1× 772  $10^{-6}$  per year in the thermal infrared window (or about 0.0003 K  $/yryr^{-1}$  change in the 773 window region); assuming the optical properties of water do not substantially change with 774 the  $\sim 0.02$  K increases seen in all the datasets considered in this paper, these very small 775 emissivity changes are of no consequence. 776

777 Land emissivity changes were estimated as follows. A global monthly mean emissivity database, the Combined ASTER and MODIS Emissivity over Land (CAMEL v003) 778 has recently been released Borbas et al. (2018). We matched the tile centers to the database 779 for the 20  $\times$  12 months spanning our 2002/09 - 2022/08 time period, and computed the 780 emissivity trends over land; the results (not shown here) were on the order of  $-1 \times 10^{-4}$ 781 and  $+3 \times 10^{-4}$  in the 800-960 cm<sup>-1</sup> and 1100-1250 cm<sup>-1</sup> regions respectively, averaged 782 over the land observations. For each tile the  $K_{emissivity}(\nu) \frac{d}{dt} \epsilon(t)$  term was estimated by 783 running SARTA with the default emissivity, then differencing with the SARTA output 784 obtained when the emissivity trends were added on. Averaged over the planet, the spec-785 tral changes arising from these emissivity changes were much smaller than the spectral 786 trends seen in Figure 3, about -0.001 K  $/year-yr^{-1}$  between 800-960 cm<sup>-1</sup> and about +0.002 787 K  $/year-yr^{-1}$  on the 1100-1250 cm<sup>-1</sup> region (which we do not use in our retrieval, since 788 many of the channels are synthetic and the real channels are drifting (Strow et al., 2021)). 789 The land only results were roughly about three times these magnitudes. Using these emis-790 sivity jacobians on the left hand side of Equation 6 and running the retrieval on the ad-791 justed spectral trends over land, resulted in about at most 0.01 K increases to the zon-792 ally averaged surface temperature changes over land; zonally averaged these largest dif-793 ferences were at about  $40^{\circ}$ N to  $60^{\circ}$ N and  $-25^{\circ}$ S to  $+15^{\circ}$ N, due to emissivity decreases; 794 the  $20^{\circ}$ N to  $+35^{\circ}$ N region which included the Sahara and swathes of Asia, had emissiv-795 ity increases but the averaged-over-land temperature decreases were small, as there were 796 offsetting emissivity increases in other land areas at the same latitudes. We did not pur-797 sue the impact of these emissivity changes further as the CAMEL database is affected 798 by the stability of the MODIS data, and our results below will not include accounting 799 for changes in land emissivity. 800

### 801 7 Results

The trends retrieved in the previous section using simulated radiance trends show 802 that the retrieval package is working as expected. Here we apply our retrieval to observed 803 AIRS L1C radiance trends and discuss the retrieved AIRS RT geophysical trends to those 804 computed directly from the ERA5/MERRA2 model fields and AIRS L3/CLIMCAPS L3 805 products. We will have an expectation that since the simulated radiance trends had no 806 noise added to them, the uncertainty in the spectral rates was lower than the actual ob-807 served spectral uncertainty; this will lead to larger uncertainties and/or errors in our re-808 trieval using observed radiance trends. 809

We will make most comparisons against NWP models and L3 products in the context of averages over the descending/night (N) and ascending/day (D) data since the MERRA2 (and GISS) datasets are only available as a D/N average; the reader is referred to the Appendix where we show a few of the D-N differences. The results are shown in the order of surface/column trends (surface temperature and column water), followed by zonal averages of the atmospheric temperature and fractional water vapor trends.

816 7.1 Skin Temperature trends

There are typically multiple (window) channels that are sensitive to a surface pressure, meaning the radiances typically have more information content for the surface tem-



**Figure 7.** Surface temperature trends dSKT/dt averaged over day and night for AIRS\_RT, and from separately fitting the monthly data in ERA5, MERRA2, AIRS L3, CLIMCAPS L3 and GISS.

perature (assuming the surface emissivity is well known and there are no clouds) rather 819 than for example air temperature. Figure 7 shows the diurnally averaged day/night (D/N)820 surface temperature trends from 6 datasets : AIRS RT, AIRS L3, CLIMCAPS L3, ERA5, 821 MERRA2 and NASA GISTEMP. AIRS RT shows an overall global warming of +0.021822  $K \neq yearyr^{-1}$ ; the cooling trends include the tropical eastern Pacific and south of Green-823 land and tropical northern Atlantic. The rest of the datasets also show similar patterns 824 of cooling in the N. Atlantic Ocean, warming over the Arctic and some degree of cool-825 ing over the Antarctic Ice Shelf/Southern Ocean as does AIRS RT. The AIRS v7 L3 826 shows some cooling over Central Africa and the Amazon not seen in the AIRS RT trends, 827 where one could expect Deep Convective Clouds and possible cloud clearing issues. We 828 also point out the AIRS L3 product has many missing values off the western coasts of 829 N. and S. America, due to cloud clearing issues. MERRA2 shows more cooling over C. 830 Africa, and just like the AIRS v7 data, a lot of cooling near the Antarctic Ice Shelf. Of 831 note here is that although CLIMCAPS uses MERRA2 as its first guess, their surface tem-832 perature trends are not similar, especially around the Antarctic where MERRA2 shows 833 strong cooling trends. Over the ocean GISS shows similar trends to what AIRS RT trends 834 show. An earlier study of Land Surface Temperatures between 2003-2017 using MODIS 835 Prakash and Norouzi (2020) shows very similar large daytime cooling trends over parts 836 of central and western Indian subcontinent that we see from our retrieval as well as di-837 rectly from the BT1231 channel trends; for tiles that straddle both ocean and land the 838 quantile method picks up the hottest observations, which especially during summer are 839 840 mostly over the Indian subcontinent. For these reasons we also have confidence in our retrieved cooling trends over for example daytime continental Central/Eastern Africa, 841 which are different from the other four day/night datasets. 842

843

The spatial correlations between AIRS RT retrieved rates and the various datasets 844 is shown in Table 2 while the cosine weighted skin temperature trends are shown in Ta-845 ble 3. By adding in the uncertainty in the trends for any of the individual models or datasets, 846 and then doing the cosine weighting, we estimate uncertainties of about  $\pm 0.015$  K /yryr<sup>-1</sup> 847 for "ALL"; the uncertainties for "OCEAN" are typically about 2/3 of that value, and for 848 "LAND" are about 4/3 of that value. We emphasize here that we use all available NWP 849 and L3 model data when computing their trends for any grid box, while the AIRS RT 850 uses only the hottest 10% of "clear" data; Strow and DeSouza-Machado (2020) showed 851 that the tropical retrieved surface temperature trends and anomalies over ocean corre-852 lated very well with those from the ERA-I Sea Surface Temperature dataset. 853

ERA5	MERRA2	AIRSL3	CLIMCAPSL3	GISS
0.72	0.59	0.80	0.89	0.77

**Table 2.** Correlations of average (nighttime,daytime) retrieved skin temperature trends fromAIRSRT, versus trends from models/products

SKT trend K $\frac{\text{yryr}^{-1}}{\text{yryr}^{-1}}$	AIRS_RT	AIRS	CLIMCAPS	ERA5	MERRA2	GISS
ALL TROPICS MIDLATS POLAR	$\begin{array}{c c} 0.020 \\ 0.011 \\ 0.029 \\ 0.032 \end{array}$	$\begin{array}{c} 0.017 \\ 0.011 \\ 0.020 \\ 0.028 \end{array}$	$\begin{array}{c} 0.021 \\ 0.012 \\ 0.028 \\ 0.033 \end{array}$	$\begin{array}{c} 0.023 \\ 0.016 \\ 0.026 \\ 0.041 \end{array}$	0.011 0.010 0.020 -0.005	$\begin{array}{c} 0.021 \\ 0.015 \\ 0.026 \\ 0.028 \end{array}$
OCEAN LAND	$ \begin{array}{c c} 0.019 \\ 0.022 \end{array} $	$0.011 \\ 0.030$	$0.019 \\ 0.024$	$0.017 \\ 0.038$	$\begin{array}{c} 0.012\\ 0.010\end{array}$	$\begin{array}{c} 0.017\\ 0.030\end{array}$

**Table 3.** Cosine weighted skin temperature trends; uncertainties are on the order of  $\pm$  0.015 K as explained in the text.

A notable outlier in this group is the MERRA2 trends, especially over land and 854 the Southern Ocean which are noticeable negative (blue) compared to the other datasets; 855 the agreement with tropical and mid-latitude oceans is much better. As noted earlier, 856 the MERRA2 monthly trends come from a combination day/night dataset that was down-857 loaded, which as seen in Figure 7 consists of trends that are both positive and negative, 858 combining to get a closer-to-zero global weighted trend. In addition MERRA2 is the only 859 one of the six that (a) does not have the extreme  $+0.15 \text{ K} \frac{\text{year-yr}^{-1}}{\text{yr}^{-1}}$  warming in the 860 northern polar region and (b) shows a lot of cooling in the Central African area. Using 861 ERA5 monthly data, we devised a test similar to the one mentioned in Section 5 to de-862 termine if the differences between MERRA2 and ERA5 surface temperature trends could 863 be due to the temporal sampling (once for MERRA2 versus eight times for ERA5). For 864 each month we matched the eight ERA5 timesteps available per month to the tile cen-865 ters and then averaged the surface temperatures per month; the ensuing geophysical time-866 series was then trended. The day/night ERA5 average of Figure 7 was compared to these 867 trends; of note are (a) we did not see the cooling in Africa and near the Antarctic that 868 is seen in MERRA2 and (b) the main differences between the 1.30 am/1.30 pm average 869 in the bottom middle (ERA5) panel were over land (all 5 continents); the histograms of 870 the differences showed the peak was typically close to 0 K /<u>yearyr</u>, but the widths over 871 land were about  $\pm 0.02 \text{K} / \text{yryr}^{-1}$  or less (compared to  $\pm 0.005 \text{ K} / \text{yryr}^{-1}$  over ocean). 872



**Figure 8.** Zonally averaged surface temperature trends for (left) sum of ocean and land point and (right) ocean only.

Both AIRS L3 and MERRA2 show cooling in the Southern Ocean; we note that although 873 MERRA2 is the *a-priori* for CLIMCAPS L3, their trends are different that those from 874 MERRA2; in fact AIRS RT shows the closest correlation to the observational CLIM-875 CAPS L3 trends. The AIRS L3 trends in the Southern Ocean region could arise because 876 of problems identifying ice during the L2 retrieval (private communication : Evan Man-877 ning (JPL) and John Blaisdell (NASA GSFC)) though the MERRA2 trends also show 878 significant cooling in that region, where few surface observations from buoys poleward 879 of 60° exist to help resolve these differences (see for example Figure 10 in Haiden et al. 880 (2018)).881

Figure 8 shows the zonally averaged total (land+ocean) and ocean only surface tem-882 perature trends. Notice how the equator to midlatitude ocean trends are almost linear 883 for all datasets, with the slope for the northern hemisphere being about double that of 884 the southern hemisphere (roughly 0.001 K  $\frac{1}{1000}$  per deg latitude). Again focusing 885 on the right hand plot, the AIRS L3 trends are negative in the Southern Ocean regions, 886 compared to the other 3 datasets, due to the cooling trends around the Antartic continent 887 shown earlier, but then agrees with most of the other datasets over the Antartic; the MERRA2 888 trends significantly differ between -90 S and -50 S. MERRA2 and ERA5 also show slightly 889 smaller warming trends in the Northern Polar, compared to the three AIRS-based datasets. 890 891

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We point out that the trends seen in Figure 7 vary noticeably at more local, regional levels and furthermore this spatial variation can differ between daytime and nighttime, evident in Figure A1 of Appendix Appendix A, and that the observational sets (AIRS\_RT, CLIMCAPS L3 and AIRS L3) had larger differences than ERA5. Discussing the possible causes of this is outside the scope of the paper.

7.2 Addition of Microwave Limb Sounder Water Vapor A-priori

The Microwave Limb Sounder (MLS), on board NASA's Aura platform, is designed for sounding of the atmosphere above 300 mb. We computed water vapor trends from the L3 data produced for that instrument (above 300 mb) and used them as an *a-priori*for the AIRS\_RTretrieval.

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Figure 9. dWVfrac/dt (left) without and (right) with MLS a-priori in the upper atmosphere

Figure 9 shows the retrieved fractional water vapor trends when the *a-priori*trend in the upper atmosphere in the left and right panels were zero, or used MLS trends, respectively. One sees that the additional information brought in by the instrument sensitive to upper troposphere humidity, significantly changes the water vapor sounding especially in the polar region by moving towards the MERRA2 and ERA5 fractional water vapor trends seen in Figure 13. We note that the results shown in this paper use the MLS *a-priori*.

### 7.3 Column water vapor trends

The left hand panel of

Column water vapor trends provide an assessment of the water vapor retrieval qual-912 ity in the lower atmosphere since this is dominated by the layers near the surface. For 913 a hyperspectral infrared sounder over ocean the 1226 (Channel ID 1511) and 1231 (Channel 914 ID 1620) spectral points are similarly impacted by surface emissivity and absorption by 915 the The water vapor information in the lowest layers is best retrieved using the weak water 916 lines in thermal infrared region. This part of the retrieval is significantly complicated by 917 the simultaneous presence of nonzero surface temperature, air temperature and water 918 vapor jacobians in this spectral region, meaning the AIRS instrument has much reduced 919 sensitivity to the water vapor continuum. However the 1226 channel is on the wing of 920 a weak water vapor line and has additional absorption from the atmospheric water vapor 921 column. Subtracting the observed brightness temperatures of these two channels BT1231 922 - BT1226 is therefore a representative approximation to (but is not equal to) the column 923 water, just as BT 1231 is a representative approximation to (but is not equal to) surface 924 temperature. For example, using the simulated AIRS L1C clearsky radiance dataset over 925 ocean we constructed for this paper using ERA5 monthly fields, we can regress the ERA5 926 column water against the brightness temperature difference to obtain  $mmw \sim 5.6$  (BT1231-BT1226) 927 + 1.0; over land the emissivity could vary rapidly enough that this approximation breaks 928 downyapor amounts in these lowest layers. In addition the changing concentration of very 929 minor gases such as CFC-11 and CFC-12 Strow and DeSouza-Machado (2020) are quite 930 evident in the spectral trends, further complicating the water vapor trend retrieval for 931 the lowest layers. 932

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Figure 10 shows the zonally averaged column water vapor trends, while the right hand panel shows the zonally averaged BT1231 - BT1226 trend (notice the multiplication



Figure 10. Zonally averaged column water vapor trends for AIRS\_RT, AIRS L3, CLIMCAPS L3, ERA5 and MERRA2.

factor of 5.6 mentioned above will roughly equalize the *y*-axis of the two panels). The gray curve is the AIRS L1C observations, while the black curve is the reconstruction from the retrieval; the rest of the curves come from the fast model simulations using the relevant model/data fields. The error bars ; not shown are the error bars which are on the order of  $\pm$  0.005 mm/year.

The column water vapor trends for , AIRS L3, CLIMCAPS L3, ERA5 and MERRA2.
 The left hand panel shows the zonal averages, while the right hand panel shows the BT1231-BT1226
 zonally averaged trends.

Close examination of the right hand panel AIRS RTis from our retrievals while 945 the rest are directly from the NWP or L3 model fields. Close examination shows the CLIM-946 CAPS L3 column water trend is nearly identical to the MERRA2 trend, as is also seen 947 in lower atmosphere water vapor trends shown later in Figure 13. Conversely the col-948 umn water vapor trends for AIRS L3 are negative in the lower troposphere in the mid-949 latitudes and tropics, which is not to be expected given that the surface temperature trends 950 are positive. AIRS RT nominally agrees with ERA5 and MERRA2 in the tropics and 951 midlatitudes, but is smaller than either in the northern polar regions. A reduced rate 952 for AIRS RT is additionally seen in the 0-50 N latitudes, where there is a larger fraction 953 of land (for which we do not use the assumption of constant relative humidity) compared 954 to the Southern Hemisphere. Screening out the tiles over land slightly improves the agreement 955 between reanalysis (ERA5, MERRA2) vs AIRS RTcolumn water trends. Examination 956 of the spectral trends in the window region does not shed any more insight into the differences, 957 as the observation spectral trends and NWP reconstructed trends are very similar and 958 we are fitting the observed trends. The magnitudes and patterns look similar to the 2005-959 2021 column water trends shown in Borger et al. (2022), which were derived using ob-960 servations from the Ozone Monitoring Instrument (OMI). We point out their 16 year zon-961 ally averaged trends look similar to the 20 year ERA5 zonally averaged column water 962 trends between  $-60^{\circ}$ S and  $-10^{\circ}$ S, but become almost a factor of 2 larger between  $-10^{\circ}$ S 963



Figure 11. The 400 mb fractional water vapor trends for (left) AIRS\_RT and (right) ERA5 show general agreement except in the Southern Polar Regions.

964	and $+$	$40^{\circ}N;$	the zonally	averaged	OMI	16	year	trends	$\operatorname{are}$	negative	in t	he p	olar	regions.	•
		-				-									

<sup>965</sup> The column water trends are summarized in Table 4.

	$\begin{array}{c} \text{DATASET} \\ \text{mm} \ \underline{/\text{year} \ \text{yr}^{-1}} \end{array}$	OMI 16 years	AIRS_RT 20 years	ERA5 20 years	MERRA2 20 years	AIRS 20 ye
~	with MLS GLOBAL (cosine average) TROPICAL	$ \begin{array}{c c} 0.051 \\ 0.083 \end{array} $	$\begin{array}{c} 0.021 \ 0.021 \\ \hline 0.028 \ 0.028 \\ \hline \end{array}$	$\begin{array}{c} 0.035\\ 0.047\end{array}$	$0.036 \\ 0.042$	-0.00 -0.01

no MLSGLOBAL (cosine average) 0.029 TROPICAL 0.039

**Table 4.** Column water trends based on OMI data (16 years) and AIRS\_RT, ERA5 and MERRA2 (20 years). The units are in mm  $/yearyr^{-1}$ ; the uncertainties are on the order of 0.1 mm  $/year yr^{-1}$  for OMI and AIRS\_RT, and half that for ERA5 and MERRA2, and AIRS L3 and CLIMCAPS L3.trends using MLS *a-priori* are shown in the table, as are trends without the MLS *a-priori* 

D/N differences (not shown) for AIRS RT were on the order of  $\pm 0.005$  mm /year 966  $yr^{-1}$  (with daytime trends being smaller over land), for AIRS L3 were on the order of 967  $\pm 0.01 \text{ mm} / \text{year} \cdot \text{yr}^{-1}$  or more (with larger values happening over the daytime tropi-968 cal oceans), while that for ERA5 and CLIMCAPS L3 were typically on the order of  $\pm$ 969  $0.03 \text{ mm} \frac{\text{year} \text{yr}^{-1}}{\text{var}}$  or less. Figure 11 shows the 400 mb fractional water vapor trends, 970 with the left panel being the AIRS RT trends while the right panel is the ERA5 trends. 971 Note that there is general agreement except in the Southern Polar region, which is as 972 also seen later in Figure 13 to some extent-in the other two observational L3 datasets 973 (AIRS v3 and CLIMCAPS). This could be related to a paper work by Boisvert et al. (2019) 974 who showed decreasing evaporation from the Southern Ocean in the 2003-2016 period 975 due to increasing ice cover. 976

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### 7.4 Zonal atmospheric temperature and water vapor trends

978 979



Figure 12. Zonally averaged dT/dt shown in 5 panels. Horizontal axis is latitude while vertical axis is pressure. The *y*-limits are between 10 to 1000 mb, on a logarithmic scale.

Figure 12 shows the zonally averaged atmospheric temperature trends from five of 980 the datasets in Figures 7,10 above. In the troposphere the AIRS RT retrievals show the 981 same general features as the trends from ERA5, though they begin to diverge in the strato-982 sphere and especially above that. In particular AIRS RT does not show warming in the 983 Southern Polar stratosphere; we have separately looked into seasonal trends and noted 984 that our retrieved September/October/November temperature trends in the upper at-985 mospheric Southern Polar regions are on the order of  $-0.12 \text{K} / \text{year} \text{yr}^{-1}$ , possibly lead-986 ing to an overall no net heating/cooling for the annual trends. In addition we point out 987 that both our results and AIRS v7 L3 show a hint of cooling over the tropical surfaces. 988 Note that CLIMCAPS is initialized by MERRA2, and their temperature trends are quite 989 similar. AIRS v7 looks similar to AIRS RT except in the tropics where it almost has 990 cooling in the lower troposphere and much more warming in the lower stratosphere. The 991 correlations between AIRS RT and the [AIRS L3, CLIMCAPS L3, MERRA2, ERA5] 992 temperature trends of Figure 12 are [0.74,0.65,0.74,0.72] respectively. 993

994

Figure 13 shows the zonally averaged atmospheric fractional water vapor trends 995  $(d/dt WV(z,t)/\langle WV(z,t) \rangle)$ . The five panels are markedly different from one another. 996 The AIRS RT trends resemble those of ERA5 in the tropical troposphere, though we 997 do not have drying in the lower tropical layers. Conversely, the observed trends in the 998 Southern Polar (AIRS L3, CLIMCAPS L3 and AIRS RT) show drying rather than wet-999 ting, though AIRS RT is less than that of CLIMCAPS/MERRA2. AIRS RT is an out-1000 lier in the upper polar atmosphere trends, as both the signals and the jacobians are close 1001 to zero. Of some concern is a little bit of drying in the northern polar region, where there 1002



Figure 13. Zonally averaged dWVfrac/dt shown in 5 panels. Horizontal axis is latitude while vertical axis is pressure. The y-limits are between 100 to 1000 mb, on a linear scale.

are low  $H_2O$  amounts leading to small jacobians. CLIMCAPS v2 looks quite similar to 1003 the MERRA2 trends. AIRSv7 shows substantial drying in the lower troposphere, and 1004 considerable wetting in the upper troposphere, compared to any of the other datasets. 1005 Spectral closure studies (using the AIRS v7  $H_2O$  trend  $\times$  the  $H_2O$  jacobians derived above 1006 from ERA5 average profiles) are not shown here, but differ noticeably from the CCR trends 1007 from AIRS v7 in the 1300-1600  $\rm cm^{-1}$  region, indicating there are inadequacies in the AIRS 1008 V7 water vapor retrievals. The correlations between AIRS RT and the [AIRS L3, CLIM-1009 CAPS L3, MERRA2, ERA5 fractional water vapor trends of Figure 13 (limited to 100 1010 mb, 1000 mb) are [0.65, 0.24, 0.36, 0.58] respectively. 1011

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### 7.5 Addition of Microwave Limb Sounder Water Vapor A-priori

1013The Microwave Limb Sounder (MLS), on board NASA's Aura platform, is designed1014for sounding of the atmosphere above 300 mb. We computed water vapor trends from1015the L3 data produced for that instrument (above 300 mb) and used them as an for the1016retrieval.

1017

dWVfrac/dt (left) without and (right) with MLS in the upper atmosphere

1018Figure 9 shows the retrieved fractional water vapor trends when the trend in the1019upper atmosphere in the left and right panels were zero, or used MLS trends, respectively.1020One sees that the additional information brought in by the instrument sensitive to upper1021troposphere humidity, significantly changes the water vapor sounding especially in the1022polar region by moving towards the MERRA2 and ERA5 fractional water vapor trends1023seen in Figure 13.

# 1024 8 Uncertainty

The uncertainties for the AIRS v7 geophysical products are impacted by radiance noise amplification due to cloud clearing Susskind et al. (2003) and the neural net first guess, while state vector errors are estimated based on regressions. CLIMCAPS L2 geophysical products are similarly impacted by cloud clearing noise in the radiances, but these are fully propagated together with geophysical error estimates from the MERRA2 first guess, through the retrieval algorithm which uses Optimal Estimation Smith and Barnet (2020). No estimate of uncertainties are available for the monthly L3 products.

The uncertainties for the AIRS RT trends is much more straightforward : the spec-1033 tral uncertainties shown in Figure 4 are used together with the state vector covariance 1034 matrices to generate the uncertainty matrix using the relevant equations of Optimal Es-1035 timation (Rodgers, 2000); we use the diagonal elements for the final uncertainties. Pan-1036 els (A) and (C) of Figure 14 shows the zonally averaged (D/N) uncertainties as a func-1037 tion of pressure and latitude. Inspection of the radiance trends uncertainties shown in 1038 the center panel of Figure 4 shows the upper atmosphere temperature sounding region 1039  $(650-700 \text{ cm}^{-1})$  has much larger uncertainty in the polar regions. The instrument and 1040 spectroscopy characteristics, coupled with these observational uncertainties, are such that 1041 for temperature the smallest errors are in the tropics while the largest errors are in po-1042 lar upper atmosphere, which are the regions below 100 mb where the ERA5 trends dif-1043 fer most from AIRS RT trends. Similarly for water vapor the larger errors are in the 1044 lower atmosphere and above about 300 mb; the constant RH assumption and MLS a-1045 *priori* help alleviate the errors. 1046

The  $h = ztest(trend, \mu = 0, trend uncertainty)Z-test$  confirmed this picture, as seen 1047 in panels (B) and (D) of Figure 414, which show the temperature and fractional water 1048 vapor trends, together with black dots marking the (latitude, altitude) points where the 1049 zero trend null hypothesis at the default significance level of trends are larger than the 1050 uncertainty in the trends, at the 5% was rejected significance level. This happens in panel 1051 (B) for the temperature trends in most of the tropical/mid-latitude free troposphere (and 1052 stratosphere) but not at the southern polar stratosphere; and in panel (D) for fractional 1053 water vapor trends in the 200-600 mb range, from the Southern Polar region to about 1054 +60 N latitude, and some spots in the Northern Polar. 1055

1056

### <sup>1057</sup> 9 Discussion

In general for surface temperature trends, the disagreements between the six sets 1058 shown in Figure 7 are over the polar regions and over land (especially over the Amazon 1059 and Central Africa) and are smallest over tropical and mid-latitude oceans, indicating 1060 the best agreements, except for slightly larger differences off the western coast of the Amer-1061 icas and Africa (which have a prevalence of MBL clouds). The atmospheric temperature 1062 trends in general agreed except for the upper atmosphere polar regions and in the high 1063 altitudes (less than about 200 mb). Similarly fractional water vapor trends differed most 1064 in the upper atmosphere (200 mb and above) and in the tropical/mid-latitude 600-800 1065 mb region. A quick glance at Figure 13 shows the former is due to lower sensitivity to 1066 upper atmosphere water vapor, leading the AIRS RT retrievals to have low values while 1067 the AIRS L2 retrieval is initialized by a neural net; conversely the latter is due to the 1068 AIRS L3 retrieval being negative while the rest were mainly positive. Similarly the AIRS RT 1069 retrieval differs above the Antarctic continent. 1070

In general the observed surface temperature trends from the AIRS\_RT retrievals agree with the ERA5 and MERRA2 trends, as well as the NASA GISS trends, except



Figure 14. Zonally averaged D/N plots of (A) temperature uncertainties in K  $\neq$ yr<sup>-1</sup> and (B) temperature trends in K  $\neq$ yr<sup>-1</sup> together with null hypothesis. (C) and (D) are the same except for fractional water vapor uncertainty and trends in 1/year. See text for more detailed explanation.

in the Southern Antarctic. That is a region where there are few surface observations; for
 retrievals there are competing effects of using ice vs ocean surface emissivity. Overall,
 the AIRS\_RT retrieved surface temperature trends are typically in between ERA5 and
 MERRA2 for land + ocean in all regimes (tropical, midlatitude and polar), though slightly
 larger overall for ocean than the two reanalysis datasets; in general they are closer to the
 ERA5 trends than the MERRA2 trends.

Strow et al. (2021) demonstrated that the long- and medium- wave channels of the 1079 AIRS instrument are radiometrically stable to better than 0.002-0.003 K /yearyr<sup>-1</sup>, which 1080 1081 is much smaller than the surface and tropospheric temperature trends in the reanalysis models, AIRS L3 data and our retrieved trends. After A separate analysis of spectral 1082 trend uncertainties after 05,10,15,20 years of observations, Figure ?? shows the trend spectral 1083 uncertainties years (not shown here) show that these uncertainties have been steadily 1084 decreasing and are now approaching this number, as can be seen in the bottom left panel 1085 of Figure 3. Furthermore, though we cannot guarantee only cloud free scenes in our cho-1086 sen Q0.90 dataset used in this paper, the high correlations between other dataset sur-1087 face trends compared to ours, is a good indication that our results come from mostly cloud-1088 free scenes, or scenes whose clouds have negligible impact on our results. 1089

The observed zonal temperature trends agree with those from the models and the 1090 AIRS L3 products, except in the polar regions. Again this could be an issue of using slightly 1091 incorrect surface emissivity for the AIRS RT retrievals. In addition we point out that 1092 since there is very little water vapor, the temperature jacobians near the surface are quite 1093 small in magnitude (compared to more humid atmospheres) and so it is difficult to sep-1094 arate out the effects of surface temperature trends versus lower atmosphere temperature 1095 and  $H_2O$  trends. The quantile construction used in this paper means that for example 1096 tiles straddling the subcontinent of India and the ocean will preferentially pick the land 1097 surface observations for daytime, which could lead to misleading trends on these coastal 1098 tiles. It is possible to subdivide the  $3^{\circ} \times 5^{\circ}$  tiles into for example  $1^{\circ} \times 1^{\circ}$  grids and do the 1099 analysis, but the number of observations per small grid cell would drop, leading to more 1100 noise in the retrieved trend. 1101

The AIRS RT retrieved absolute column water trends are equal to/slightly larger 1102 than ERA5/MERRA2 in the tropics and below both of them in the midlatitudes; AIRS RT 1103 ocean column water trends were slightly smaller than both ERA5 and MERRA2 over 1104 ocean, and in-between them over land. We note the difficulties we have retrieving  $H_2O$ 1105 close to the surface and in the upper atmosphere. This is simply a consequence of the 1106 sensitivity of the infrared sounder, namely most of the averaging kernels peak in the 300-1107 600 mb range. AIRS RT column water trends agree with those from ERA5 and MERRA2 1108 column water trends in the tropics; nevertheless even with expected lowered sensitivity 1109 to water vapor in the lower altitudes, we were able to retrieve similar column water va-1110 por trends to the NWP models both in the tropics and in the mid-latitudes. The dif-1111 ferences become more acute in the polar regions since the low average amounts of wa-1112 ter vapor mean the water vapor jacobians are very small, as were the observed trends 1113 in the WV channels. However, we point out that our column water trends, which are both 1114 quite sensitive to water vapor in the lower atmosphere, are in good agreement with those 1115 from NWP models. 1116

We point our here that our results are relatively robust to changes in the covariance or Tikonov parameter settings. For instance changing them by factors of two would keep the trends about the same, though of course the uncertainties would change.

Given the complex numerical algorithms used in both the reanalysis models and the AIRS L3 retrievals as well as those in the AIRS\_RT trends, it is difficult to offer precise explanations for any of the trends shown above. There are however a few general points that can be made. The first is that since infrared instruments are sensitive to the 300-800 mb region and lose sensitivity outside this, the retrievals from AIRS\_RT and

AIRS L3 have difficulties with water vapor in the lower (Planetary Boundary Layer) and 1125 upper troposphere/lower stratosphere. One way to mitigate this is to use trended data 1126 from external sources ; for in the *a*-priori, while keeping the *a*-prioritrends for all other 1127 parameters as 0. For example we have shown we can use the MLS data above 300 mb 1128 without significantly degrading the AIRS RT retrieval in the middle and lower atmo-1129 sphere; conversely the CLIMCAPS retrievals are initialized by MERRA2 and while they 1130 can pull out weather signals, their L3 trends are still quite closely tied to the MERRA2 1131 trends. The tropical and mid-latitude ocean surface temperature trends from the numer-1132 ical models that assimilate data, L3 products and AIRS RT are very similar; however 1133 they start to show differences where there are few *in-situ* data combined with problems 1134 with ice identification (surface emissivity)/cold temperatures which exacerbate the drift-1135 ing AIRS detector problems Strow et al. (2021), such as the Arctic and Southern Ocean. 1136

## 1137 10 Conclusions

We have described a novel method to obtain global thermodynamic atmospheric 1138 climate trends, starting from infrared allsky hyperspectral observations which are then 1139 subset for "nominally clear" scenes. Our retrieved trends are derived using trends from 1140 well characterized (radiometrically stable) radiances and from zero *a-priori* (except for 1141 a constant relative humidity assumption). This makes them much more direct and trace-1142 able than trends from traditional L2 retrieval algorithms, which use complicated *a-priori* 1143 information. We also did "radiative closure" tests by running the monthly NWP or L3 1144 fields through a radiative transfer model to compare the spectral trends so obtained against 1145 the observed spectral trends, which showed the most disagreement in the water vapor 1146 sounding regions. 1147

The temperature and water vapor trends retrieved from the "nominally clear" ra-1148 diance trends resemble those computed from monthly ERA5 and MERRA2 reanalysis. 1149 The radiative spectral closure helps identify the cause of differences in the geophysical 1150 trends, rather than solely attributing them to deficiencies (eg the well known reduced 1151 sensitivity to water vapor near the boundary layer and above 200 mb) with our retrieval. 1152 For example the AIRS RT temperature trends are quite similar to the reanalysis (MERRA2/ERA5) 1153 trends, while the water vapor (and/or Relative Humidity) trends are quite different, es-1154 pecially in the lower troposphere and upper troposphere, which is clearly manifest as dif-1155 ferences in the spectral trends in the water vapor sounding region. 1156

The 20 years of AIRS observations were binned into nominal  $3 \times 5$  degree grid boxes 1157 covering the planet, with a time step of 16 days, from which anomalies and trends were 1158 obtained. To alleviate the reduced sensitivity of hyperspectral sounders to water vapor 1159 in the lower atmosphere we used an assumption of 0.01 increase in relative humidity to 1160 initialize the *a-priori* lower atmosphere fractional water vapor rates, while we similarly 1161 used Microwave Limb Sounder trends as an *a-priori* to address the high altitude water 1162 vapor deficiencies caused by lower sensitivity to upper atmosphere water vapor. New or 1163 updated time dependent surface emissivity databases may become available in the fu-1164 ture, enabling us to include those effects into Equation 4. Problems in the polar regions 1165 and Planetary Boundary Layer water vapor retrievals will be harder to overcome since 1166 there is very little sensitivity to water vapor in these regions, together with fewer obser-1167 vations to compare against, though more work is planned to address both of these. 1168

In this paper we used the 90th quantile (Q0.90) nominally "hottest" observed BT1231 data to form a time series over which to obtain radiance trends, after establishing that the spectral trends from this quantile differed by less than about  $\pm 0.0015 \text{ K} / \text{yryr}^{-1}$ from the 50th (or average) quantile. In the future we plan to base the data subset selection on MODIS cloud products (obtained at 1 km resolution compared to the AIRS 15 km resolution). In any case the AIRS L1C Q0.90 spectral trends used for the AIRS\_RT results are very comparable to trends from quality assured binned AIRS CCR data Manning

(2022). The quantile method allows us to select which data to use in the trends : we have 1176 explored doing the trend retrievals using the cloud fields contained in ERA5, together 1177 with the TwoSlab cloud algorithm De Souza-Machado et al. (2018) to compute jacobians 1178 1179 when clouds are present, together with trends from the Q0.50 dataset described above. The retrieved geophysical trends resemble those described above in the mid to upper at-1180 mosphere, and differ in the lower atmosphere, but more work is needed on this and is 1181 not discussed further. Longwave clear sky flux trends (both outgoing top-of-atmosphere 1182 and incoming bottom-of-atmosphere) and climate feedbacks will be discussed in a sep-1183 arate paper. 1184

While the Aqua platform is scheduled to be terminated within the next few years, 1185 copies of near identical CrIS instruments are already in orbit, and more will be launched 1186 over the next few years, till at least 2040. The Climate Hyperspectral Infrared Radiance 1187 Product (CHIRP) Strow et al. (2021) will seamlessly combine the AIRS data between 1188 2002-2015 to CrIS data from 2015-2040 to obtain a 40 year observational radiance record 1189 over which to study climate. This availability means that AIRS RT and future AIRS/CrIS 1190 versions, is well positioned to enable climate analysis of geophysical trends for years to 1191 come. 1192

# <sup>1193</sup> Appendix A Day versus Night surface temperature trend differences

Figure A1 shows the (top) daytime and (middle) nighttime surface temperature 1194 trends; from left to right the datasets are (observational) AIRS RT, AIRS L3, CLIM-1195 CAPS L3 and (reanalysis) ERA5. In general the AIRS observational datasets show en-1196 hanced daytime cooling over the Indian subcontinent and Central Africa, compared to 1197 the ERA5 model; they also show daytime warming trends over continental Europe and 1198 central Asia and the Amazon are larger than during the nightime. With the large ocean 1199 heat capacity and smaller land heat capacity, the land is expected to show more of a di-1200 urnal cycle than ocean. ERA5 sees warming over Eastern/Central Africa during daytime 1201 while the observational datasets see cooling. Similarly the three observational datasets 1202 see more daytime cooling over the Indian sub-continent and south eastern Australia than 1203 does ERA5; we omit more detailed analysis in this paper. During the nighttime, the AIRS 1204 L3 product has cooling over C. Africa and parts of the Amazon. The day-night differ-1205 ences are seen in the bottom row of the same figure. Note the colorbar is the same for all three rows. The differences are close to zero over the ocean. AIRS RT and CLIM-1207 CAPS L3 see more daytime cooling over E. Africa and the Indian subcontinent. Over-1208 all the magnitude of the day - night differences for the observations are larger for the AIRS 1209 observational datasets than for ERA5. ERA5 also sees negative differences over Central 1210 Asia compared to the AIRS observational datasets, which see positive differences (higher 1211 surface temperature trends during the daytime). 1212

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The atmospheric temperature and fractional water vapor day-night differences are quite small (compared to the average values) and not shown here; AIRS L3 shows noticeable more wetting of the 600-800 mb region during daytime versus nightime, compared to the other three.

# 1218 Data availabilityOpen Research Section

The AIRS L3 and CLIMCAPS L3 data products, as well as the AIRS L1C radiances are freely available to the public on the NASA servers. MERRA2 and ERA5 and GISTEMP model output are also freely available.



**Figure A1.** Top two rows : The (top) day and (middle) night surface temperature trends for AIRS RT, AIRS L3, CLIMCAPS L3 and ERA5. Third row (bottom) is the D-N difference.

### 1222 Author contribution

Sergio DeSouza-Machado prepared the manuscript with contributions from all co-authors,
 and did most of the data analysis in this manuscript. L. Strow envisaged the concept
 of tiling the AIRS observations into tiles and drove the research work. R. Kramer provided
 valuable advice regarding the methods and data analysis, as well in preparing the manuscript.

# 1228 Competing interests

1229

The authors declare that they have no conflict of interest.

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